

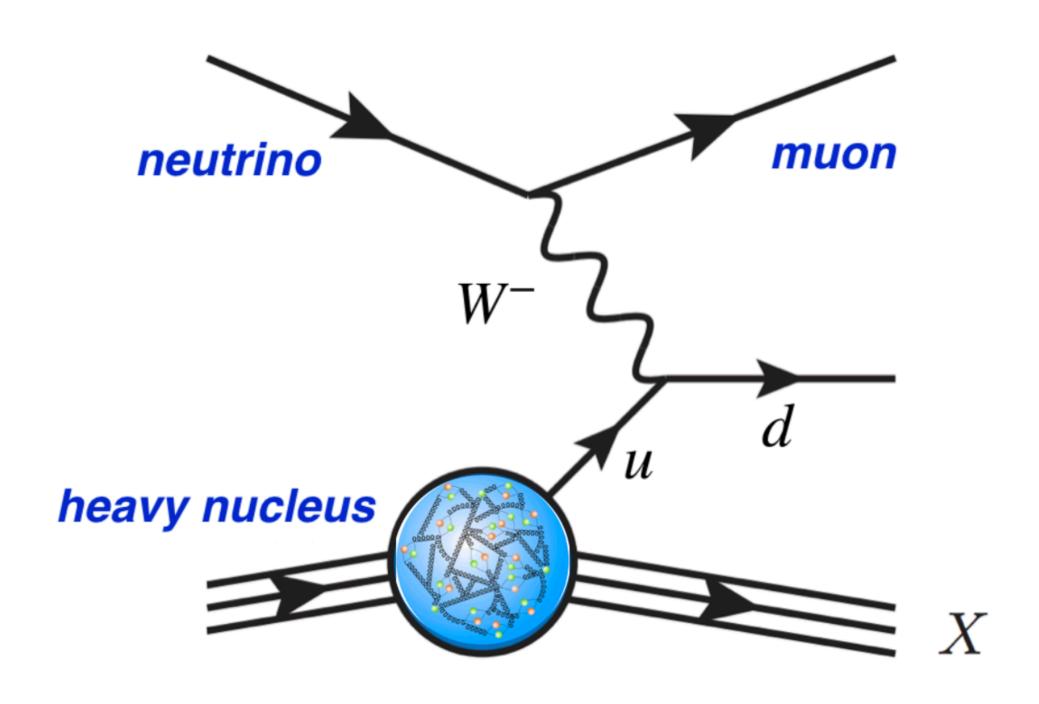


What can we learn from high-energy neutrino-nucleon collisions?

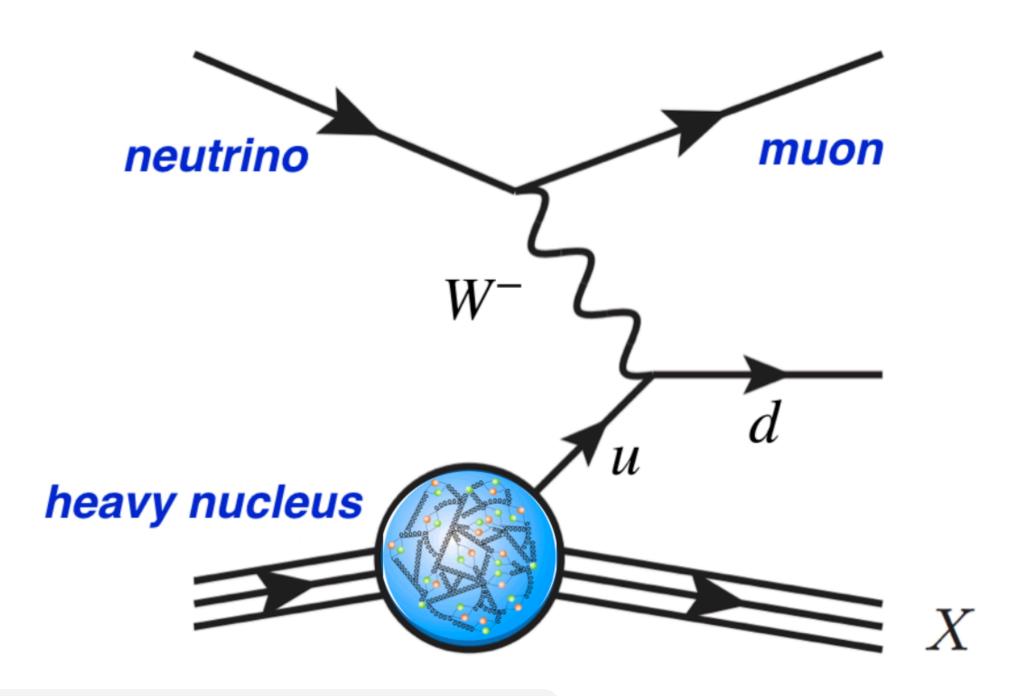
Juan Rojo

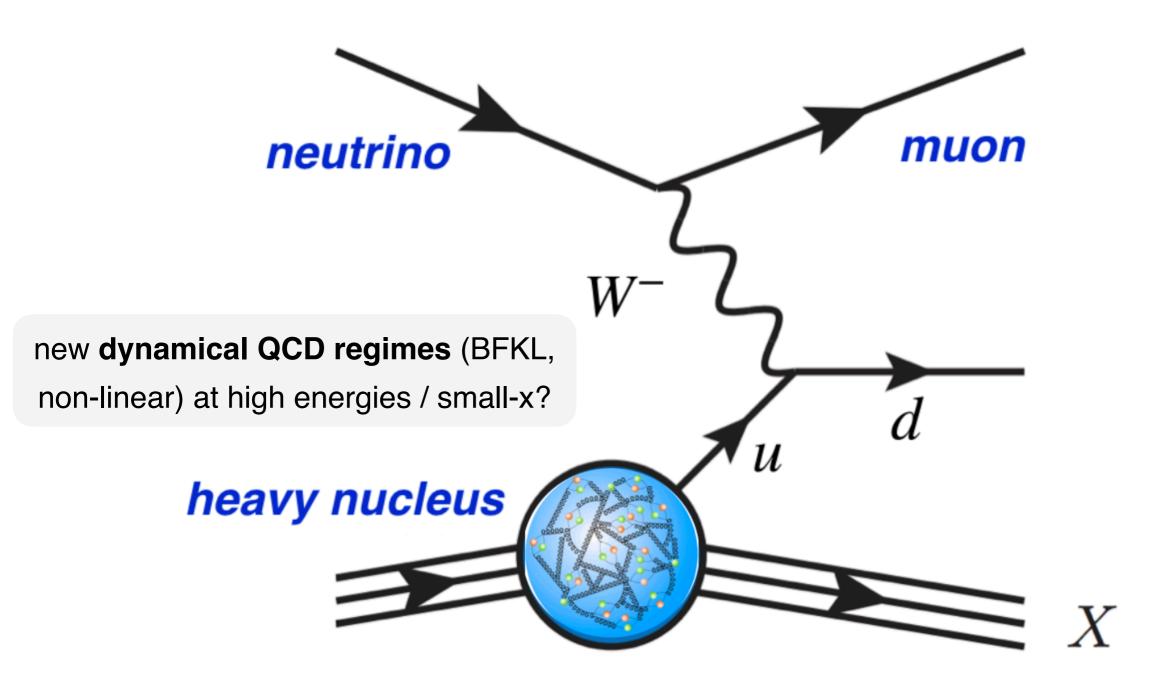
VU Amsterdam & Theory group, Nikhef

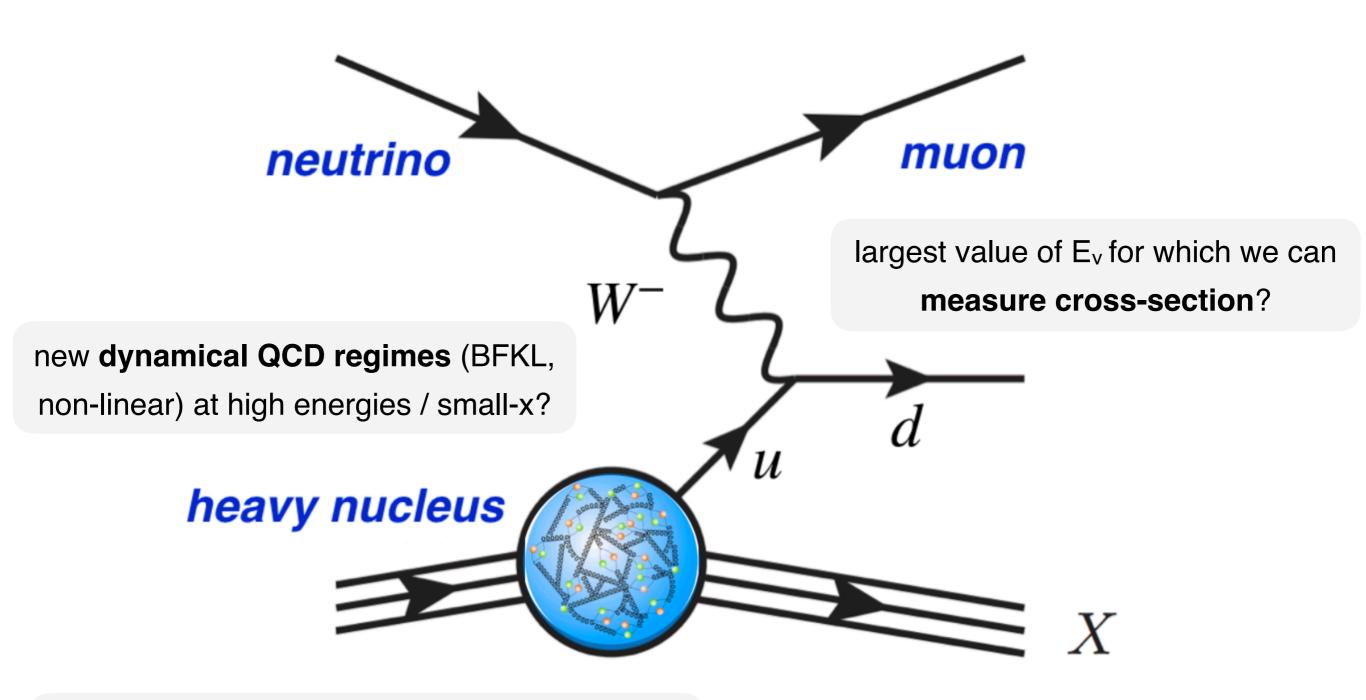
Nikhef Workshop Theory Meets Experiment
High-energetic neutrino scattering processes and
interactions of cosmic rays

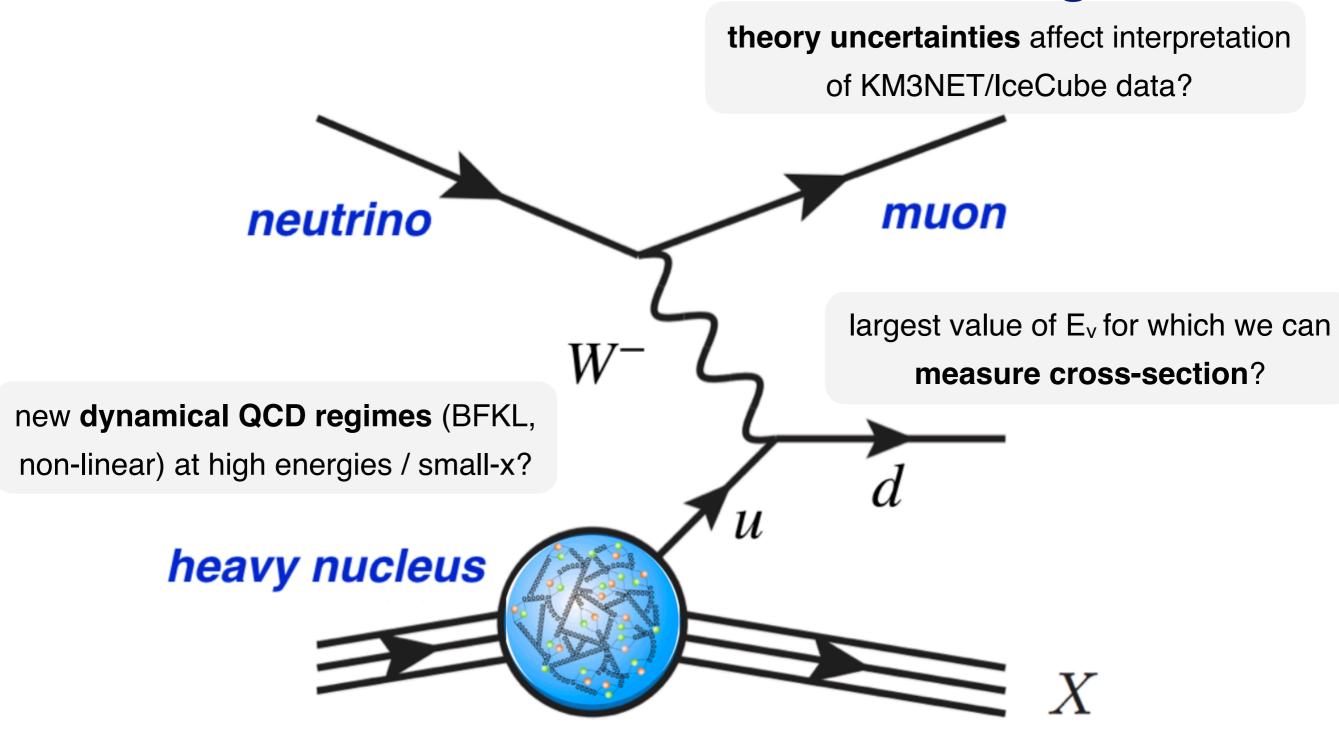


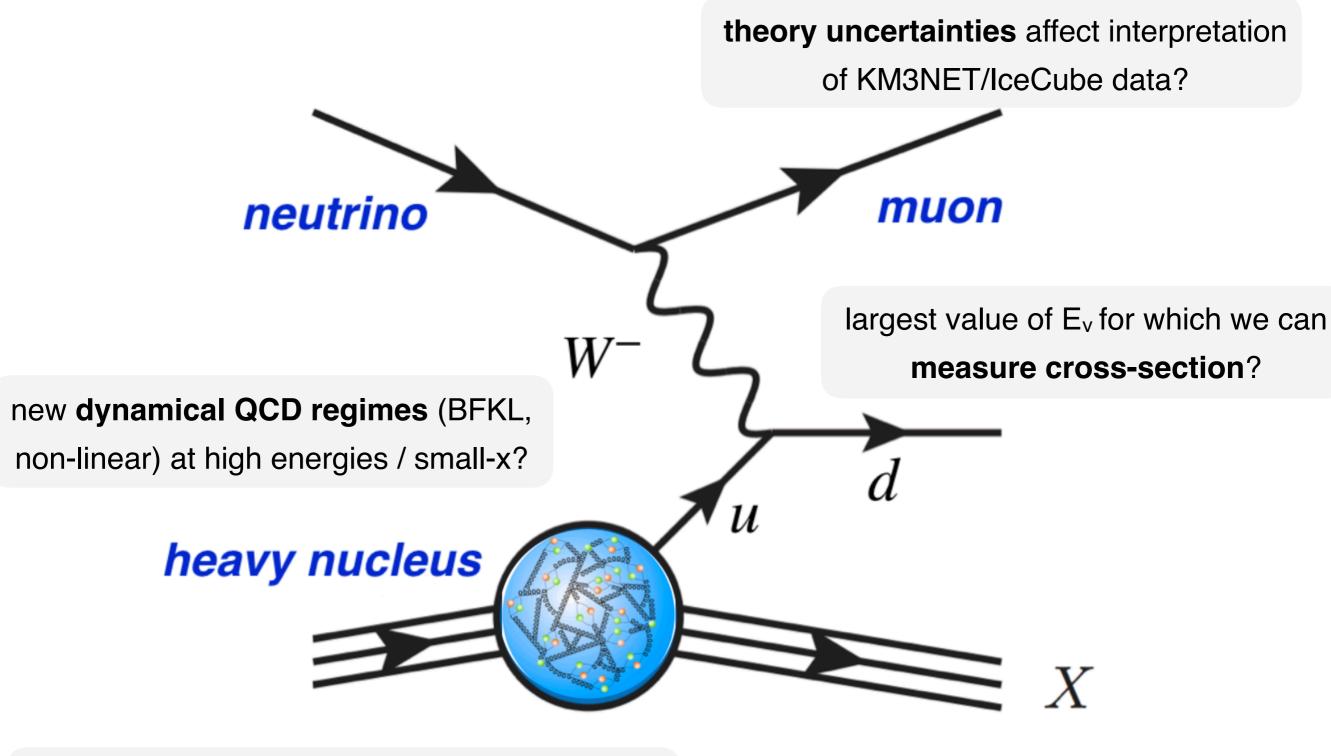
Rhorry's talk: review of the theory of neutrino-nucleus scattering





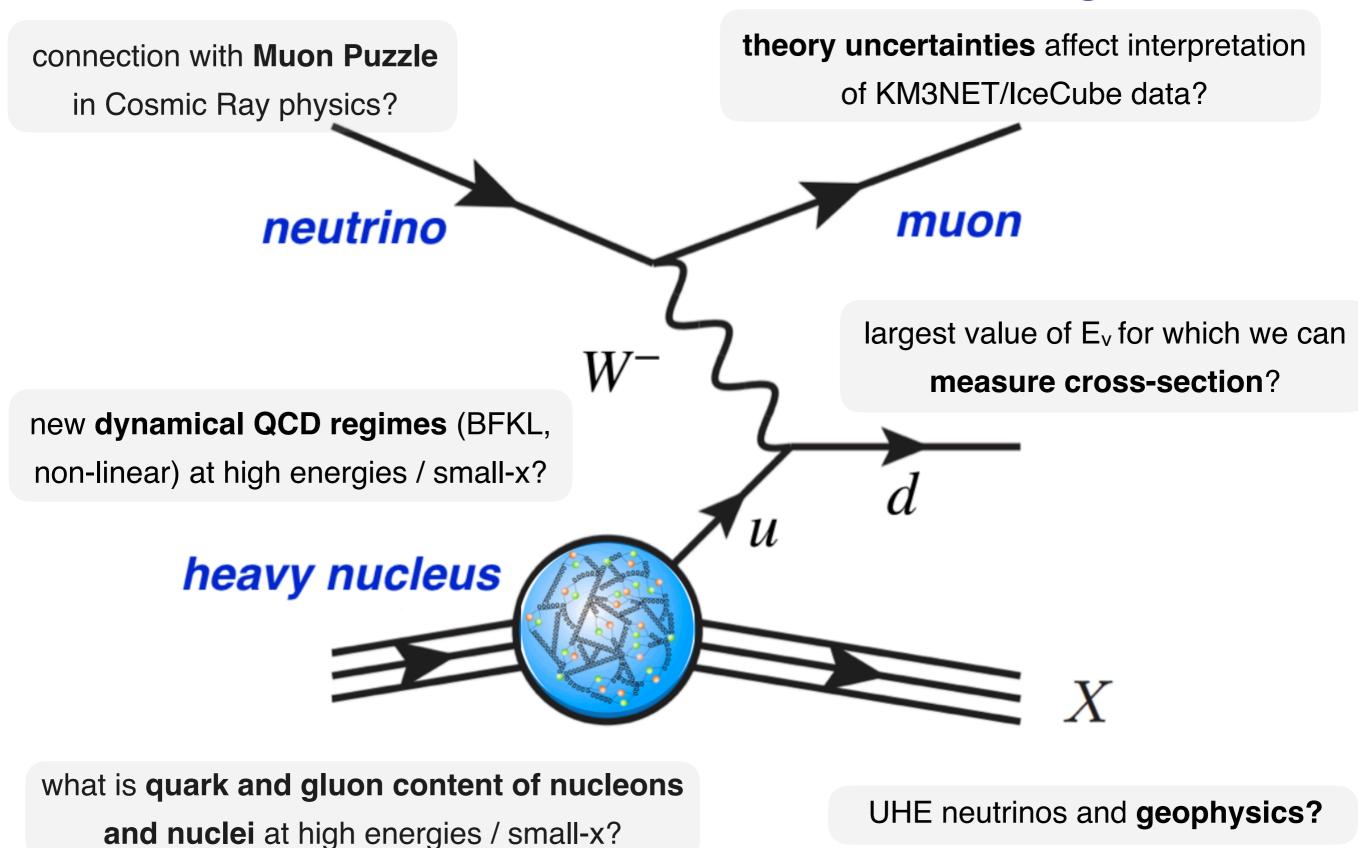






what is quark and gluon content of nucleons and nuclei at high energies / small-x?

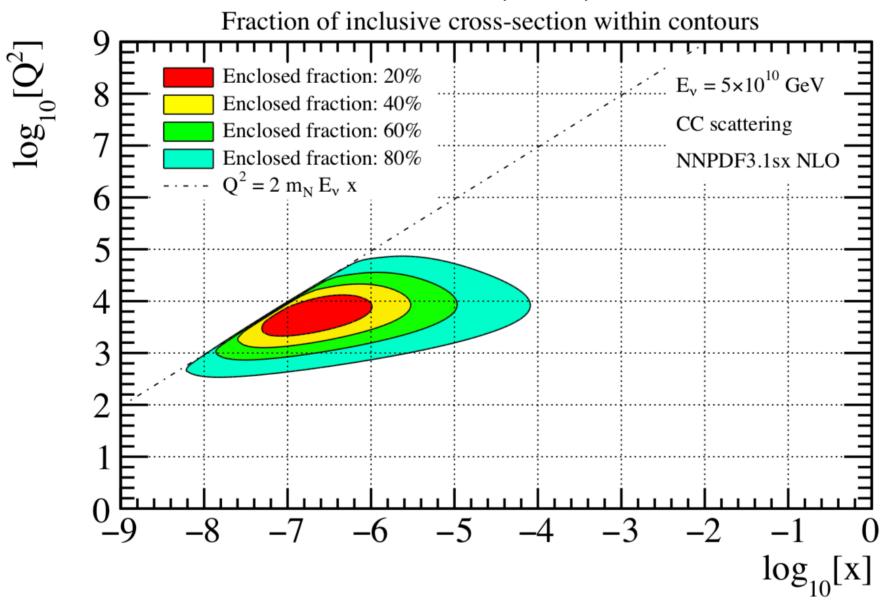
UHE neutrinos and **geophysics**



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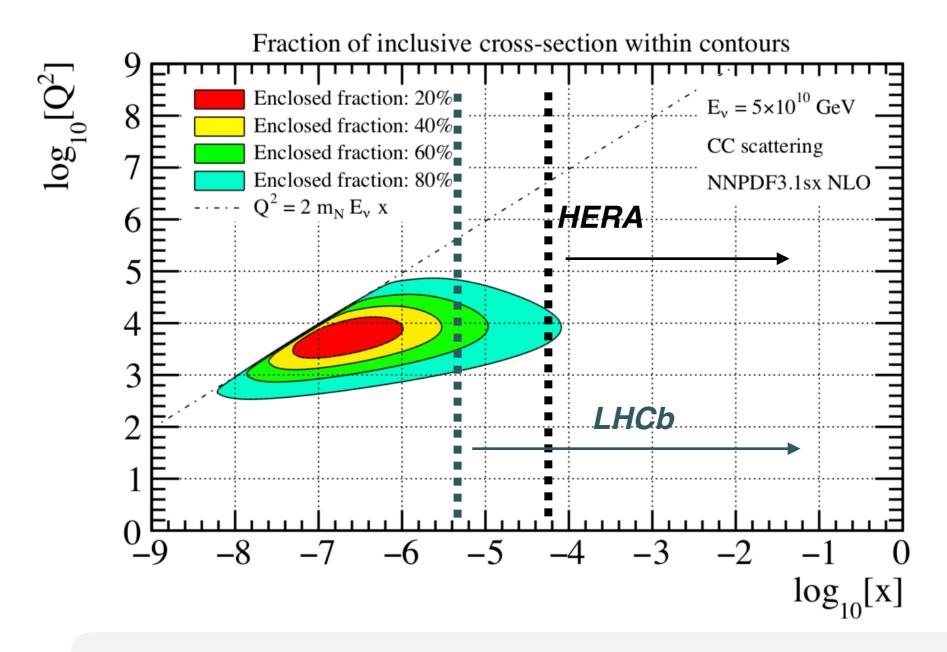
Neutrino telescopes as probes of small-x QCD





sensitive to **small-x quarks** (and thus gluons via evolution) down to $\mathbf{x} \approx \mathbf{10^{-8}}$ and $\mathbf{Q} \approx \mathbf{M_W}$

Neutrino telescopes as probes of small-x QCD

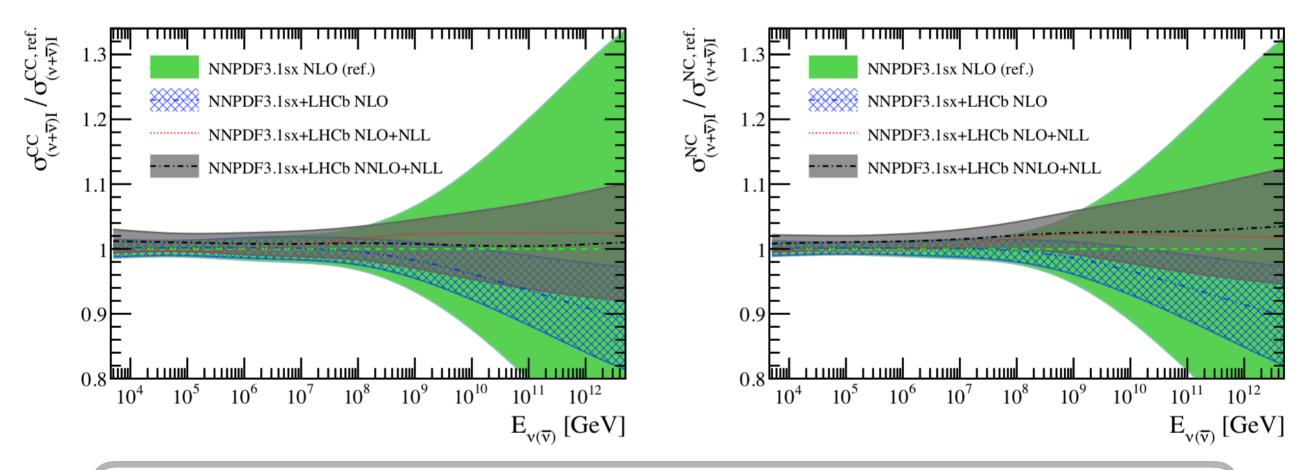


sensitive to **small-**x **quarks** (and thus gluons via evolution) down to $x \approx 10^{-8}$ and $Q \approx M_W$

unique access to small-x in the foreseeable future!

Neutrino telescopes as probes of small-x QCD

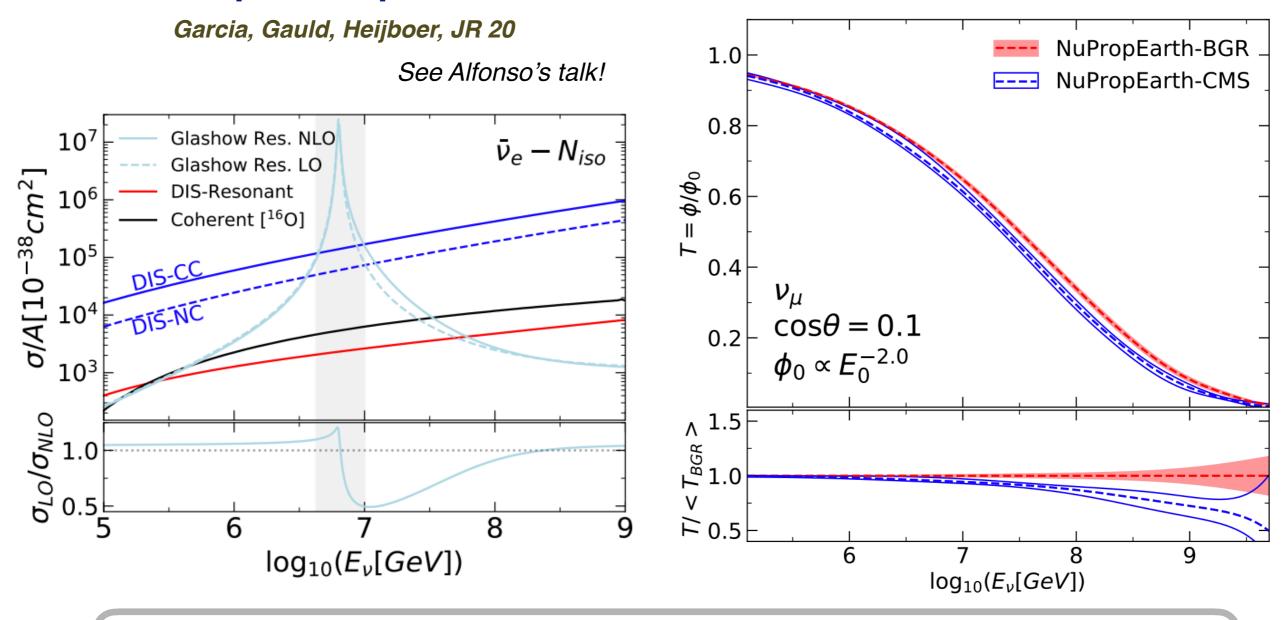
Bertone, Gauld, JR 18



State-of-the-art predictions for ultra-high energy neutrino interactions

- BFKL small-x effects in PDFs and deep-inelastic structure functions
- Constraints on small-x PDFs from LHCb charm production
- Accounting for **nuclear corrections** and heavy-quark-initiated contributions

Complete predictions for attenuation rates



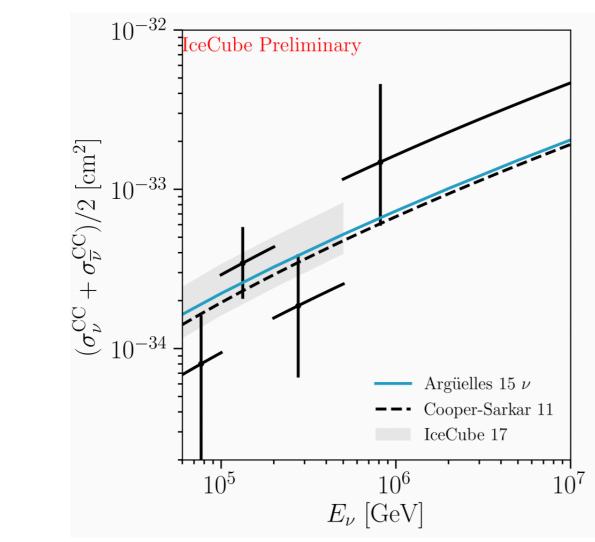
State-of-the-art predictions for ultra-high energy neutrino attenuation rates

- All subleading channels included, implemented in GENIE Monte Carlo
- Used to constrain nuclear structure, Earth model: geophysics with neutrinos

Cross-section measurement: how far in energy?

charged-current scattering at very high-energies:

fundamental measurement of neutrino telescopes, impossible in other facilities



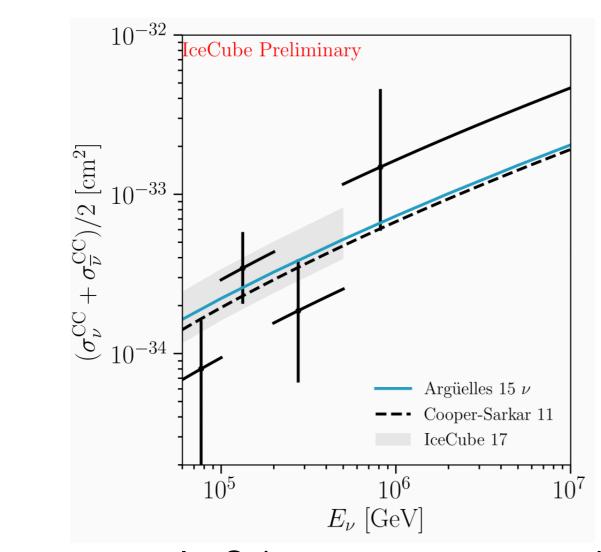
current IceCube measurements sensitive to

$$\sqrt{s} = \sqrt{2m_N E_{\nu}} \simeq 2 \,\mathrm{TeV}$$

Cross-section measurement: how far in energy?

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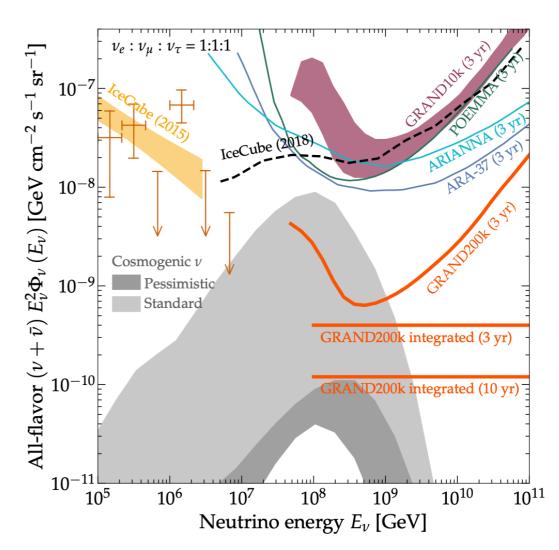
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Juan Rojo



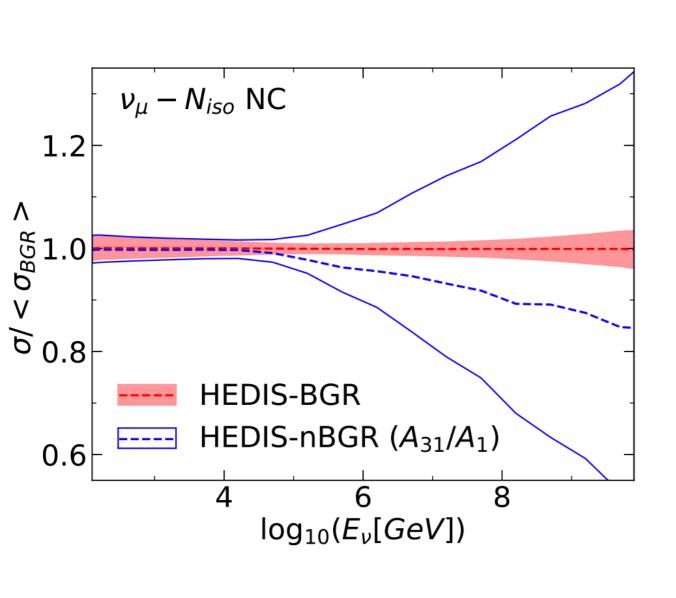
GRAND could reach

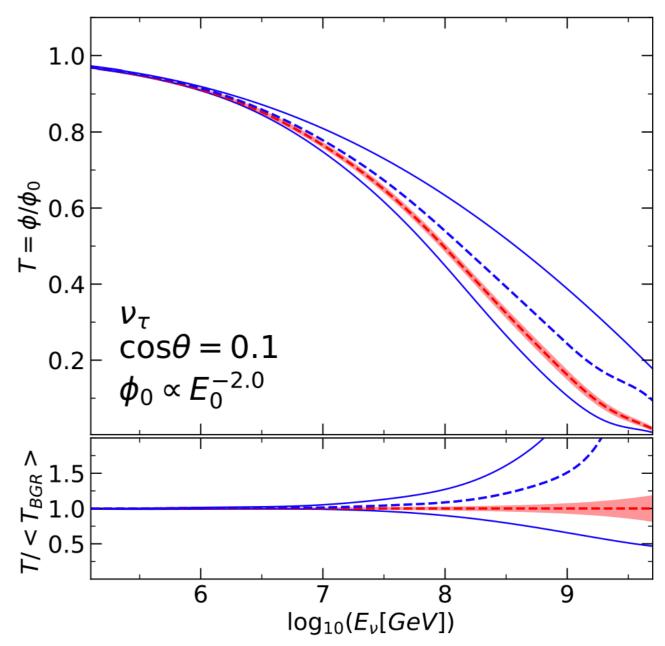
$$\sqrt{s} = \sqrt{2m_N E_{\nu}} \simeq 100 \,\mathrm{TeV}$$

what would be precision of such measurement?

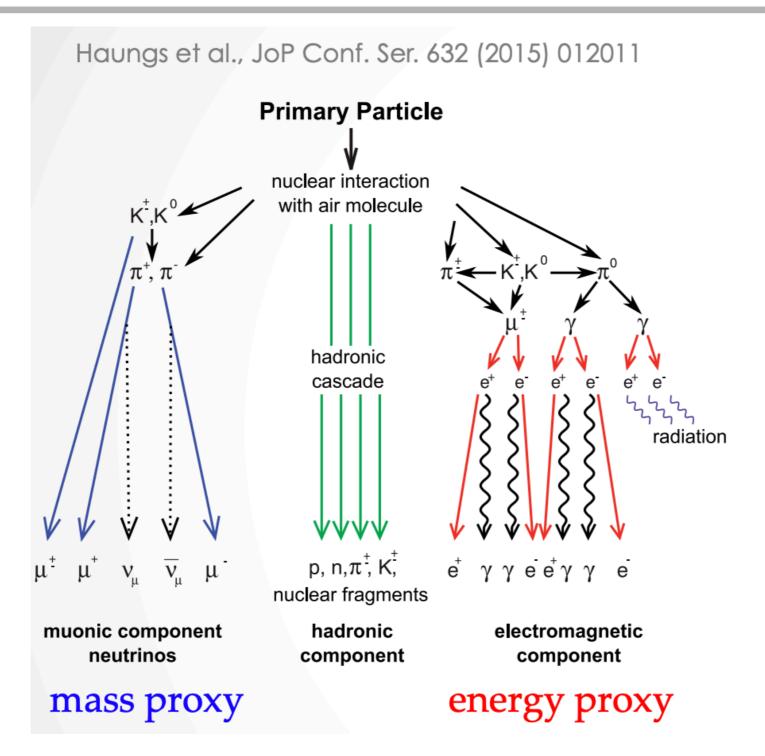
Nuclear effects in neutrino DIS

nuclear modifications (shadowing) main source of TH error at high-energies exploit information from **LHC pPb collisions** (WIP) + future colliders (eg **EIC**)

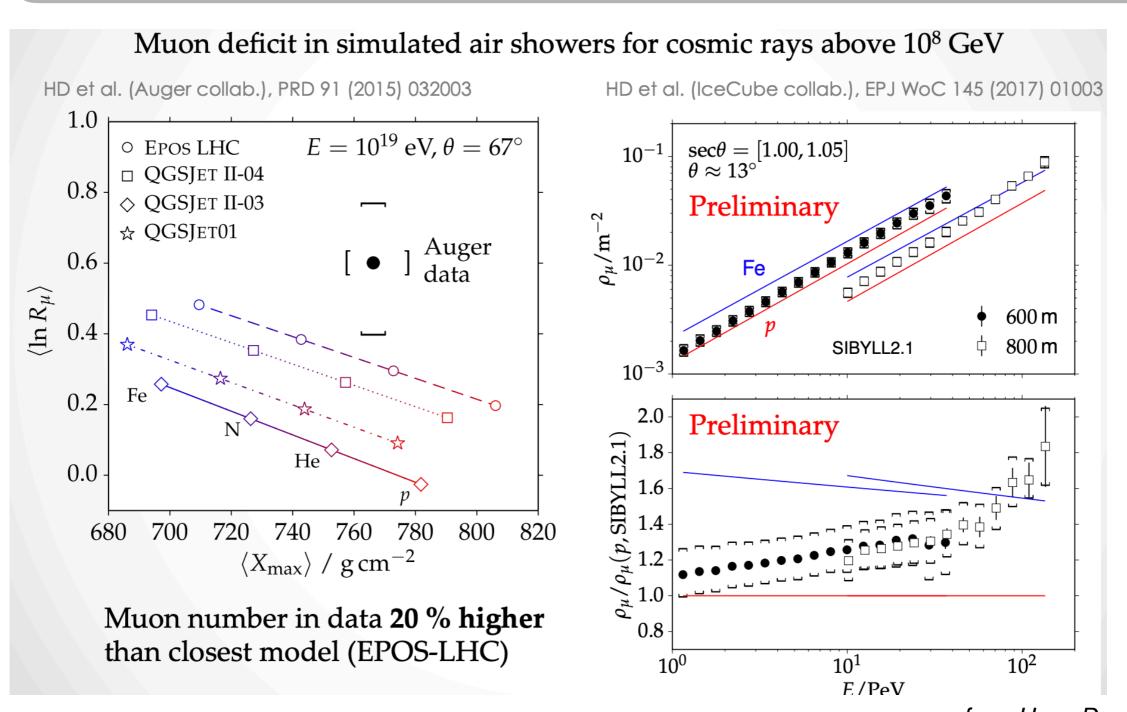




cosmic ray collisions: highest energy natural collider available! properties of cosmic ray primary extracted from resulting shower

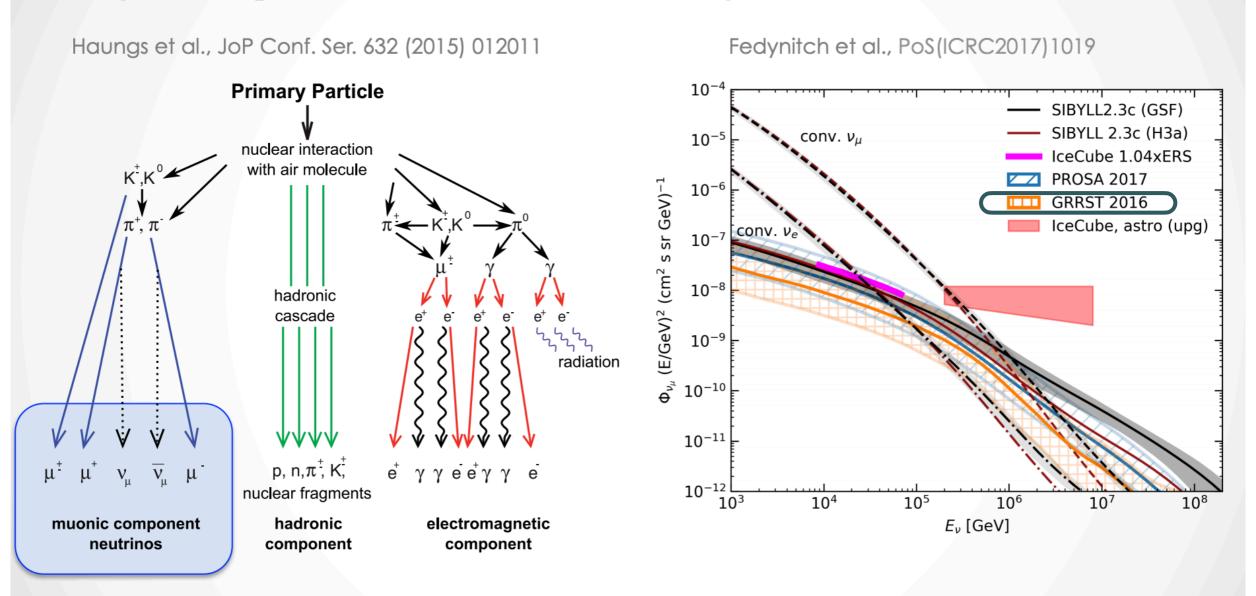


the number of muons in energetic air showers **systematically higher** than theory simulations: where is the problem?



The same energetic cosmic rays responsible for the main backgrounds to astrophysical neutrinos: **prompt neutrinos from charm decay**

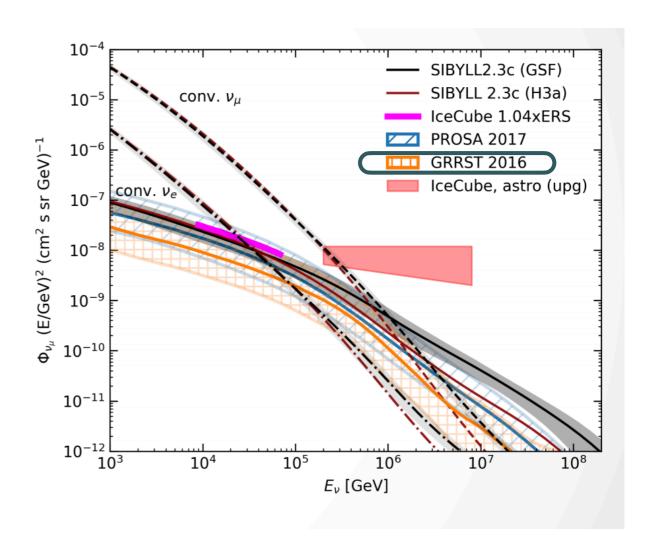
To get better predictions of atm. neutrino background for neutrino observatories



Muon production and **neutrino** production in air showers **coupled** Wrong number of muons = wrong prediction of atm. neutrino background

maybe muons from **heavy quark decays** are being mismodeled? how we can use best the **constraints from the LHC** to solve the muon puzzle can we, QCD theorists, apply our neutrino know-how to **cosmic ray physics?**

Is there interest to kick-start collaborations between Nikhef CR experimentalists and theorists?



For discussion

UHE neutrino and cosmic ray physics are exciting fields with lots of potential for cross-field collaborations and synergies: **from LHC to astroparticles!**

Some possible topics to further explore in the discussion:

precise pQCD predictions for charm production in cosmic rays?

Connection with models of nuclear structure and **pPb collisions at LHCb**?

ultimate precision for the **measurement** of the UHE neutrino cross-sections?

Where else is the input from theorists required by **KM3NET/Auger?**