



# QCD and neutrino physics at the FPF and the role of neutrino interaction modelling

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FLArE Far Forward Physics working group, 07.03.2023

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TeV neutrino deep-inelastic scattering: charged-current counterpart of Electron Ion Collider





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Collider counterpart of high-energy cosmic rays interactions, including prompt neutrino flux



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the feasibility of neutrino physics at the LHC already demonstrated by FaserNu and SND@LHC experiments



# The Forward Physics Facility

A new facility in a tailor-made underground cavern hosting a suite of **far-forward experiments** suitable to detect **long-lived BSM particles** and **neutrinos** produced at the HL-LHC



no modifications to the HL-LHC infrastructure or operations required

# FPF physics potential

#### ☑ BSM searches

- Light BSM particles produced in the very forward direction
- Decaying dark sector long-lived particles (dark photons, dark Higgs, heavy neutral leptons...)
- Milli-charged particles, dark matter scattering, ...
- Meutrino physics
  - Tau neutrino studies (10k tau neutrino interactions, current world sample <20)</p>
  - Separation of tau neutrino / anti-neutrino, constrain tau neutrino EDM
  - Tau neutrino decays into heavy flavour (connection with LHCb LFV anomalies)
  - EFT constraints on neutrino interactions

**OCD**, hadron structure, and astroparticle physics

- Neutrino cross section measurements (energy region not covered by any other experiment)
- Neutrino DIS to constrain proton and nuclear structure
- Testing BFKL dynamics in LHC collisions, modelling charm, hadron production in forward region
- Key input for neutrino (IceCube, KM3NET) and cosmic ray astroparticle experiments



# QCD and Neutrino Interactions at the FPF

#### Neutrino production and detection



#### Neutrino production and detection



#### Neutrino production and detection







10

1.5

2.5

2

 $\log_{10}(E_v)$ 

3

[GeV]

3.5

FPF unique probe of intrinsic heavy quarks, key feedback for prompt neutrino flux calculations at IceCube/KM3NET





- Excellent complementarity between EIC (neutral current) and FPF (charged current) measurements of DIS structure function on proton and nuclear targets
- A joint analysis of EIC+FPF data markedly improves the (n)PDF reach of individual experiments







- At the FPF the flux and flavour of the incoming neutrinos depends on the energy: we can either take it from existing calculation or constrain it from the data
- Focus on charged-current inclusive scattering, with a single charged lepton in final state. Extend to semi-inclusive processes (e.g. dimuon production) afterwards
- Model how each experiment measures final-state particles to reconstruct the DIS kinematics

Assume that we can access the outgoing charged lepton energy, the lepton scattering angle, and the total hadronic energy or invariant mass of the hadronic final state

$$\left(E_{\ell},\ heta,\ W^2
ight)$$
 or  $\left(E_{\ell},\ heta,\ E_h
ight)$ 

Final sector of the sector of

$$\left(x,\ Q^2,\ E_{
u}
ight)$$
 or  $\left(x,\ Q^2,\ y
ight)$ 

by using the following equations



nb ideally we'd like to over-constrain the kinematics by measuring more variables than unknowns

Given the DIS kinematics of an event, the interaction probability will be proportional to the doubledifferential DIS cross-section

$$\begin{aligned} \frac{d^2 \sigma^{\nu A}(x,Q^2,y)}{dxdy} &= \frac{G_F^2 s/2\pi}{\left(1+Q^2/m_W^2\right)^2} \left[ (1-y) F_2^{\nu A}(x,Q^2) + y^2 x F_1^{\nu A}(x,Q^2) + y \left(1-\frac{y}{2}\right) x F_3^{\nu A}(x,Q^2) \right] \\ \frac{d^2 \sigma^{\nu A}(x,Q^2,y)}{dxdy} &= \frac{G_F^2 s/4\pi}{\left(1+Q^2/m_W^2\right)^2} \left[ Y_+ F_2^{\nu A}(x,Q^2) - y^2 F_L^{\nu A}(x,Q^2) + Y_- x F_3^{\nu A}(x,Q^2) \right] \end{aligned}$$

- Traditionally neutrino measurements are presented at the level of individual structure functions, but this requires extra assumptions: cleaner to measure directly the reduced cross-section
- Free number of events in a given bin will be given by

$$N_{\text{ev}}(x \in [x_{\min}, x_{\max}], Q^2 \in [Q^2_{\min}, Q^2_{\max}, E_{\nu} \in [E_{\nu,\min}, E_{\nu,\max}]) \propto \int_{x_{\min}}^{x_{\max}} dx \int_{Q^2_{\min}}^{Q^2_{\max}} dQ^2 \int_{E_{\nu,\min}}^{E_{\nu,\max}} dE_{\nu} \frac{d^2\sigma(x, Q^2, E_{\nu})}{dxdy} f(E_{\nu})$$

$$experiment-dependent factor scattering incoming cross-section neutrino flux$$

We have produced dedicated sets of FPF pseudo-data for each of the proposed experiments and for different systematic correlation models: used as input in NNPDF and xFitter PDF projections



Large statistics for inclusive DIS, with uncertainties at the few % level for most of the bins

- When added to PDF4LHC21 via PDF profiling, large impact specially in **up and down valence PDFs**
- Full FPF statistics crucial: PDF impact much reduced e.g. for FASERnu pseudo-data
- Work in progress: implement realistic systematic correlation models



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	lepton energy E	lepton angle θ	charged lepton sign	hadronic final state
FaserNu2	E <sub>l</sub> > 100 GeV δE <sub>l</sub> = 30%	tan(θ) < 0.5 δθ = 1 mrad	Yes, for muons	$E_h$ accessible, charm ID possible, $\delta E_h = 30-50\%$
AdvSND@LHC	E <sub>I</sub> > 20 GeV (muon)	θ < 0.15 rad (muon) θ < 0.5 rad (electron, tau)	Yes	E <sub>h</sub> accessible
FLArE	E <sub>I</sub> < 1 TeV, δE <sub>I</sub> = 5% (electron) E <sub>I</sub> < 2 GeV (muon)	$\theta < 0.5 \text{ rad, } \delta\theta =$ 15 mrad (electron) $\theta < 0.4 \text{ rad (muon)}$	Maybe, for muons	$E_h$ accessible, $\delta E_h = 30\%$

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# Neutrino Structure Functions from GeV to EeV energies

arXiv:2302.08527

# The role of the low-Q region

FPF neutrinos have **energy distributions** dominated by region [**100 GeV**, **5 TeV**]. How reliably can we predict their cross-sections and event rates?

23





### The role of the low-Q region

FPF neutrinos have **energy distributions** dominated by region [100 GeV, 5 TeV]. How reliably can we predict their cross-sections and event rates?



$$\sigma(\mathbf{E}_{\nu}) = \int_{Q_{\min}^2}^{2m_N \mathbf{E}_{\nu}} dQ^2 \left[ \int_{Q^2/(2m_N y \mathbf{E}_{\nu})}^1 dx \, \frac{d^2 \sigma}{dx dQ^2}(x, Q^2, \mathbf{E}_{\nu}) \right]$$

#### State of the art

The Bodek-Yang model is popular to describe inelastic neutrino DIS structure functions

based on **effective leading-order PDFs** (GRV98LO) supplemented to phenomenological scaling variables and *K*-factors to improve agreement with data

$$f_i^{\text{LO}}(x,Q^2) \to f_i^{\text{LO,BY}}(\xi,Q^2) \qquad \xi = \frac{2x(Q^2 + m_f^2 + B)}{Q^2 \left[1 + \sqrt{1 + (2m_N x)^2/Q^2}\right] + 2Ax}$$

**Limitations** of the BY model of neutrino structure functions:

- Obsolete PDF parametrisation that ignores constraints from the last 25 years
- Neglects higher-order QCD corrections (can be up to 100%)
- Does not provide **uncertainty estimate**, difficult to assess its accuracy and precision
- Cannot be systematically improvable e.g. by new data

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# The NNSFv approach

Use available data on neutrino-nucleus scattering to parametrise and determine inelastic structure functions by means of the NNPDF fitting methodology



- Finily  $\Im$  This data-driven parametrisation is made to **converge to the pQCD calculation** for large enough  $Q^2$  values as implemented with Lagrange multipliers
- $\frac{1}{2}$  In the neutrino energy region sensitive only to Q > few GeV, replace by pQCD calculation

consistent determination of neutrino structure functions valid for 12 orders of magnitude from  $E_{nu} = few \text{ GeV}$  up to  $E_{nu}=10^{12} \text{ GeV}$ 

## The NNSFv results



Smooth matching between data-driven and pQCD regions, uncertainty estimate in whole energy range

- Structure functions and integrated cross-sections available via user-friendly LHAPDF grids
- For the first time, a **unique theory prediction** for neutrino inelastic scattering suitable for neutrinos with energies from a few GeV up to the multi-EeV region

#### The NNSFv results



#### The NNSFv results





- Good agreement with available neutrino structure function and cross-section data
- Robust estimate of all relevant sources of experimental and theory uncertainties
- Model-independent determination of nuclear corrections to free-nucleon scattering

## **NNSFv cross-sections for FPF simulations**



Peliable state-of-the-art predictions for differential neutrino cross-sections at FPF energies

#### Using NNSFv for neutrino simulations

- Free NNSFV structure functions are provided in terms of fast LHAPDF interpolation grids
- Frey can be readily used in **GENIE** by means of the **HEDIS** package (official GENIE release)
- Same GENIE/HEDIS interface: access other cross-section models like Bodek-Yang and BGR18
- Implementation in other neutrino event generators straightforward: no reason not to adopt NNSFv in your neutrino scattering simulations!



(Z, A) [target]	low- $Q$ grid	high- $Q$ grid
(1,2) [D]	NNSFnu_D_lowQ	NNSFnu_D_highQ
(2,4) [He]	NNSFnu_He_lowQ	NNSFnu_He_highQ
(3,6) [Li]	NNSFnu_Li_lowQ	NNSFnu Li highQ
(4,9) [Be]	NNSFnu_Be_lowQ	NNSFnu_Be_highQ
(6, 12) [C]	NNSFnu_C_lowQ	NNSFnu_C_highQ
(7, 14) [N]	NNSFnu_N_lowQ	NNSFnu_N_highQ
(8, 16) [O]	NNSFnu_O_lowQ	NNSFnu_O_highQ
(13, 27) [Al]	NNSFnu_Al_lowQ	NNSFnu_Al_highQ
(15, 31) [Ea]	NNSFnu_Ea_lowQ	NNSFnu_Ea_highQ
(20, 40) [Ca]	NNSFnu_Ca_lowQ	NNSFnu_Ca_highQ
(26, 56) [Fe]	NNSFnu_Fe_lowQ	NNSFnu_Fe_highQ
(29, 64) [Cu]	NNSFnu_Cu_lowQ	NNSFnu_Cu_highQ
(47, 108) [Ag]	NNSFnu_Ag_lowQ	NNSFnu_Ag_highQ
(50, 119) [Sn]	NNSFnu_Sn_lowQ	NNSFnu_Sn_highQ
(54, 131) [Xe]	NNSFnu_Xe_lowQ	NNSFnu_Xe_highQ
(74, 184) [W]	NNSFnu_W_lowQ	NNSFnu_W_highQ
(79, 197) [Au]	NNSFnu_Au_lowQ	NNSFnu_Au_highQ
(82, 208) [Pb]	NNSFnu_Pb_lowQ	NNSFnu_Pb_highQ

# The HEDIS package

- Lead developer: Alfonso Garcia Soto (MIT & IFIC)
- Original goal was to extend coverage of GENIE to neutrino energies above 1 TeV
- Current implementation, when combined with NNSFv, allows calculations of inelastic scattering for all energies from a few GeV to the multi-EeV regime
- Current status of GENIE in the high energy regime:
  - DIS based on Bodek-Yang model -> optimised for low Q<sup>2</sup>.
  - Structure Function =  $C_{ij}$  LO  $\otimes$  PDF LO (GRV98 Q<sup>2</sup>[0.8,2.10<sup>6</sup>]).
  - Contributions from heavy quarks are not included.



#### **Targeted experiments**

- KM3NeT:
  - Already using GENIE (both DIS and HEDIS) in its simulation framework (gSeaGen).
- IceCube(-Gen2):
  - Uses HEDIS as an auxiliary tool to crosscheck their simulation framework at HE.
- Neutrino facilities at LHC:
  - Data in 2021-2023 (FASERnu).
  - Overlapping region between DIS & HEDIS (~0.1-1TeV).
  - Not sure what simulation package they are using so far.
- Others: GVD-Baikal, P-ONE, GRAND, etc.
- New extension allows UHE interaction -> HEDIS
  - Newer PDFs with broader Q<sup>2</sup> phase space.
  - Structure Functions =  $C_{ij}$  NLO $\otimes$ PDF NLO.
  - Account for the heavy quark contributions.

Alfonso Garcia Soto, GENIE Users Meeting, Dec 2021

#### The HEDIS package

- $T = \phi(E)/\phi_0(E_0)$ measured **φ(E)** neutrino flux detector Successfully deployed for many phenomenological neutrino applications **Earth** Attenuation rates of UHE neutrinos Forecasts for UHE cross-section measurement  $\phi_0(E_0) \propto E_0^{-\gamma}$ incoming neutrino flux Center-of-mass energy  $\sqrt{s}$  [GeV] 1.0 $10^{0}$  $10^{1}$  $10^{2}$  $10^{3}$  $10^{4}$  $10^{5}$ Neutrino-nucleon cross section,  $\sigma_{NN}^{
  m CC}$   $[10^{-38}~{
  m cm}^2]$ GGM-SPS 8 LEP Tevatron LHC  $v_{\tau}$ 0.9  $10^6$  $\cos\theta = 0.01$ φ/φ Φ/φ  $\phi_0 \propto E_0^{-2.0}$  $10^{5}$ SciBooNE 1 II CCFR 97  $10^{4}$ CDHS 87 NuPropEarth-BGR (ALLM) 0.7 NuPropEarth-CMS (ALLM)  $10^{3}$ NuPropEarth-CMS (BS) Accelerator  $\nu$ Accel.  $\nu$ Ultra-high-energy cosmic  $\nu$ High-energy cosmic  $\nu$  $10^{2}$ IceCube-Gen2 Radio (10 yr, projected) IceCube tracks (avg.  $\nu + \bar{\nu}$ ) TauRunner-CMS (ALLM-CMS) 0.6 (IceCube 17) Using cosmogenic  $\nu$  flux. IceCube showers (avg.  $\nu + \bar{\nu}$ fit to TA UHECRs (Bergman & van Vliet) NuTauSim-CTW (ALLM) 10<sup>1</sup> (Bustamante & Connolly 17 Using IceCube  $\nu_{\mu}$  flux (9.5 yr), extrapolated to UHE IceCube HESE (avg.  $\nu + \bar{\nu}$ ) 0 Using cosmogenic flux (IceCube 20) 1.1 from all AGN (Rodrigues et al.) 100 ٨ 1.0 <sup>BGR</sup> FASERi  $\nu N$  DIS prediction b.f.  $\pm 1\sigma$  (BGR18) 10  $10^{10}$  $10^{-1}$  $10^{0}$  $10^{1}$  $10^{2}$  $10^{3}$  $10^{4}$  $10^{11}$  $10^{5}$  $10^{6}$  $10^{8}$  $10^{2}$  $10^9$ V Neutrino energy,  $E_{\nu}$  [GeV] 2004.04756 F 0 6 8 9 2204.04237  $\log_{10}(E_{\nu}[GeV])$ 
  - And ready to become the go-to tool for the **modelling of neutrino interactions in the FPF era**
  - Next milestone: exclusive event generation for high-energy neutrino-nucleus ĕ interactions at (N)NLO QCD matched to parton showers!

## Summary and outlook

- The FPF would realise an exciting program in a broad range of topics from BSM and long-lived particles to neutrinos, QCD, and hadron structure, with connections to astroparticle physics
- Fine FPF would continue the long tradition of neutrino DIS @ CERN now with TeV beams
- High-energy neutrino DIS would open a new probe to proton and nuclear structure: a chargedcurrent counterpart of the Electron Ion Collider
- Ongoing studies demonstrate the quantitative reach of FPF measurements to constrain proton and nuclear structure, in particular quark & antiquark flavour separation
- Precision QCD and neutrino physics at the FPF requires state-of-the-art modelling of neutrino inelastic structure functions and cross-sections,
- The NNSFv calculation of neutrino structure functions and cross-sections, valid for all energies of phenomenological relevance, is available via the HEDIS module of GENIE