

Introduction to Elementary Particles (TN2811)
Theory Lecture 2

Dr Juan Rojo

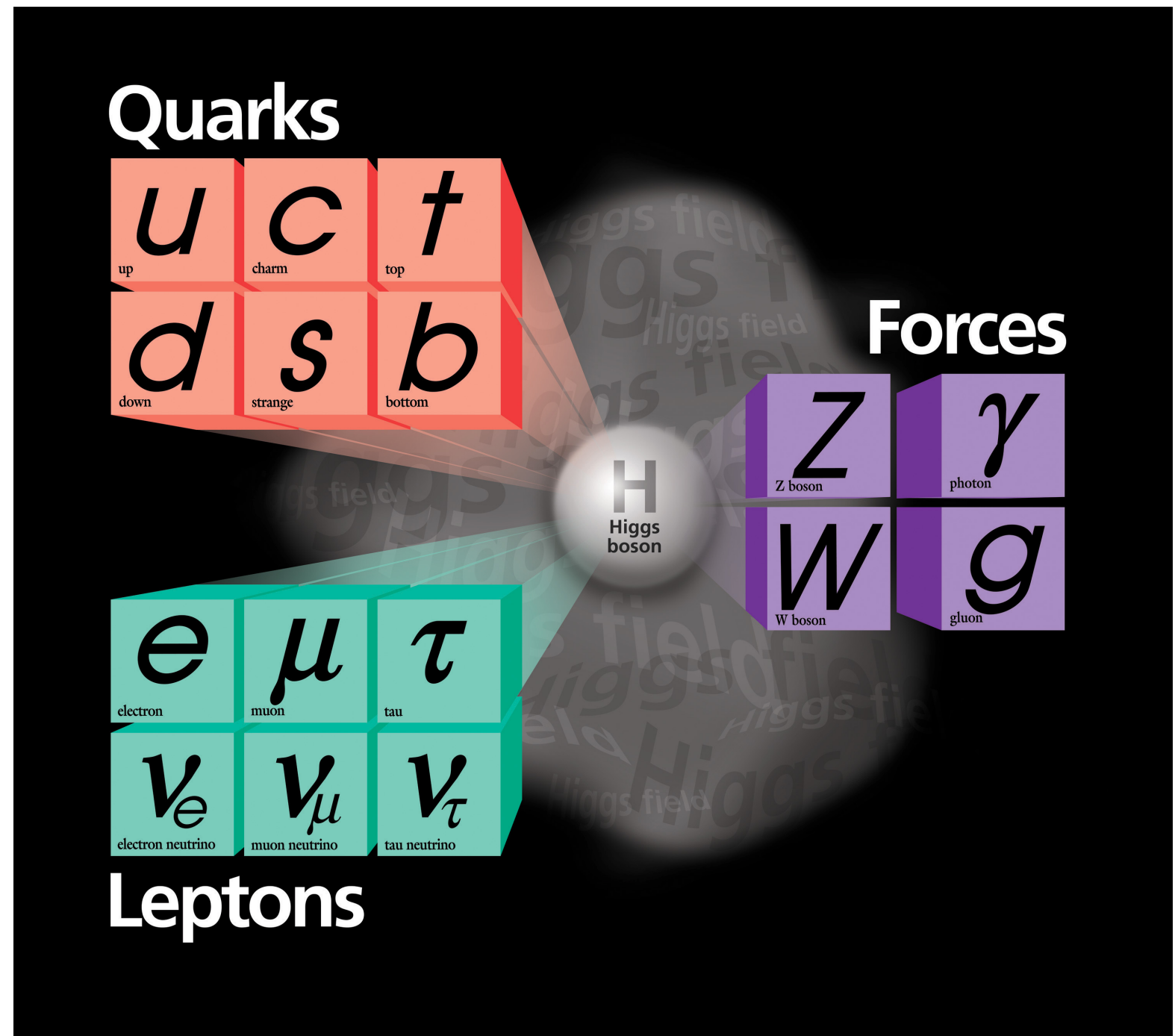
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Recap: the Standard Model

- ✓ The **Standard Model (SM)** of particle physics explains a wide variety of microscopic phenomena in a unified framework: **Quantum Field Theory**
- ✓ **Matter content** composed by **six quarks** and **six leptons**, organised in **three families**
- ✓ Interactions **between matter particles** are governed by **gauge bosons**: **photons** (electromagnetism), **W and Z bosons** (weak force), and **gluons** (strong interaction)
- ✓ The last ingredient is the **Higgs Boson**, provides mechanism by which **particles acquire mass**



Today's lecture

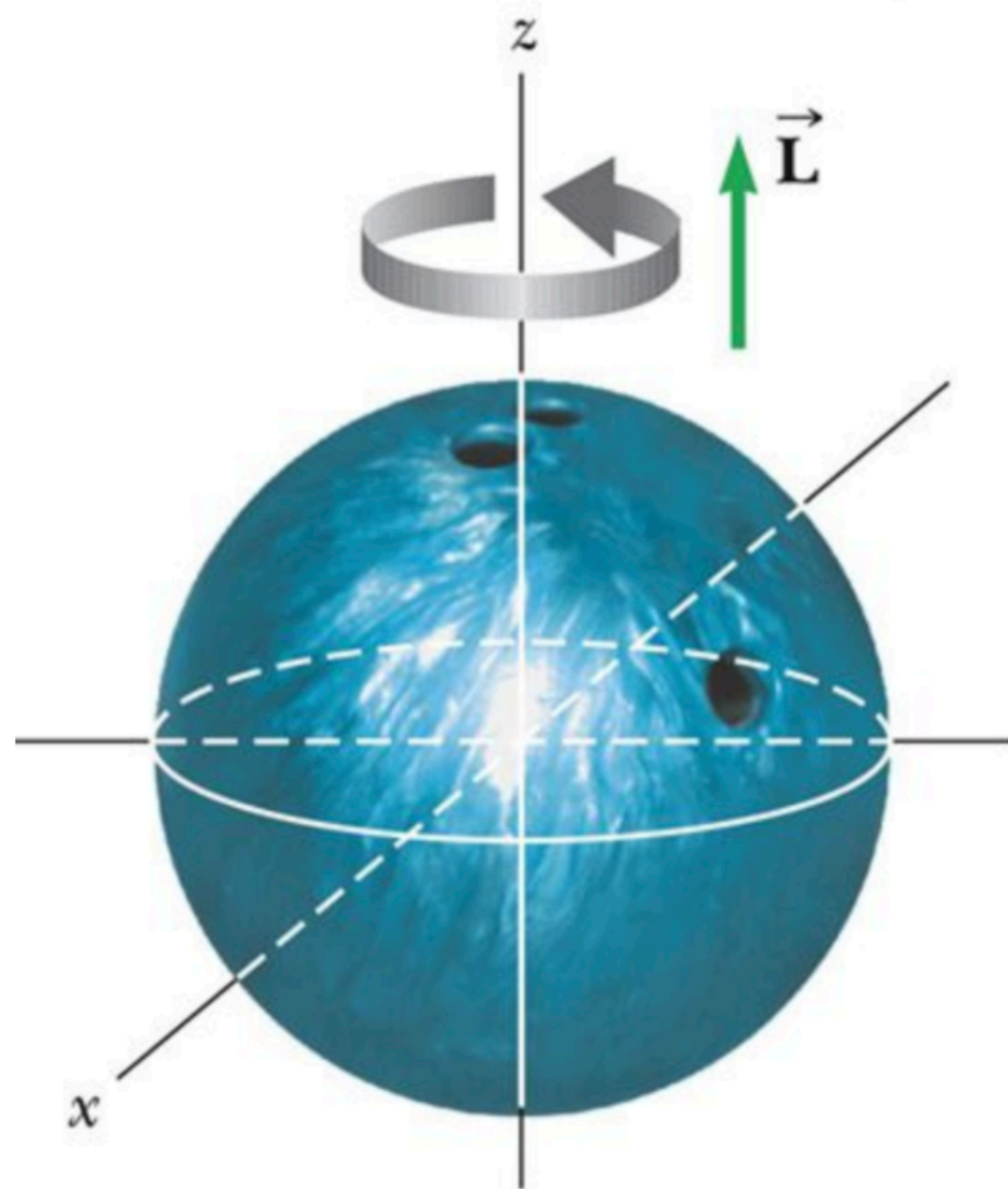
- ☑ Spin and statistics: fermions and bosons
- ☑ More about neutrinos: **how large a neutrino detector** needs to be?
- ☑ The gravitational interaction, black holes, and **gravitational wave detection**
- ☑ Elements for **Special Relativity** relevant for Elementary Particle Physics

Spin in quantum theory

Intrinsic angular momentum

In classical mechanics, a solid body rotating around one axis has associated **angular momentum**, which is **conserved** in the absence of external forces

Bowling ball (classical mechanics)



$$|\vec{L}| = I \times \omega = \left(\frac{2}{5} MR^2 \right) \omega$$

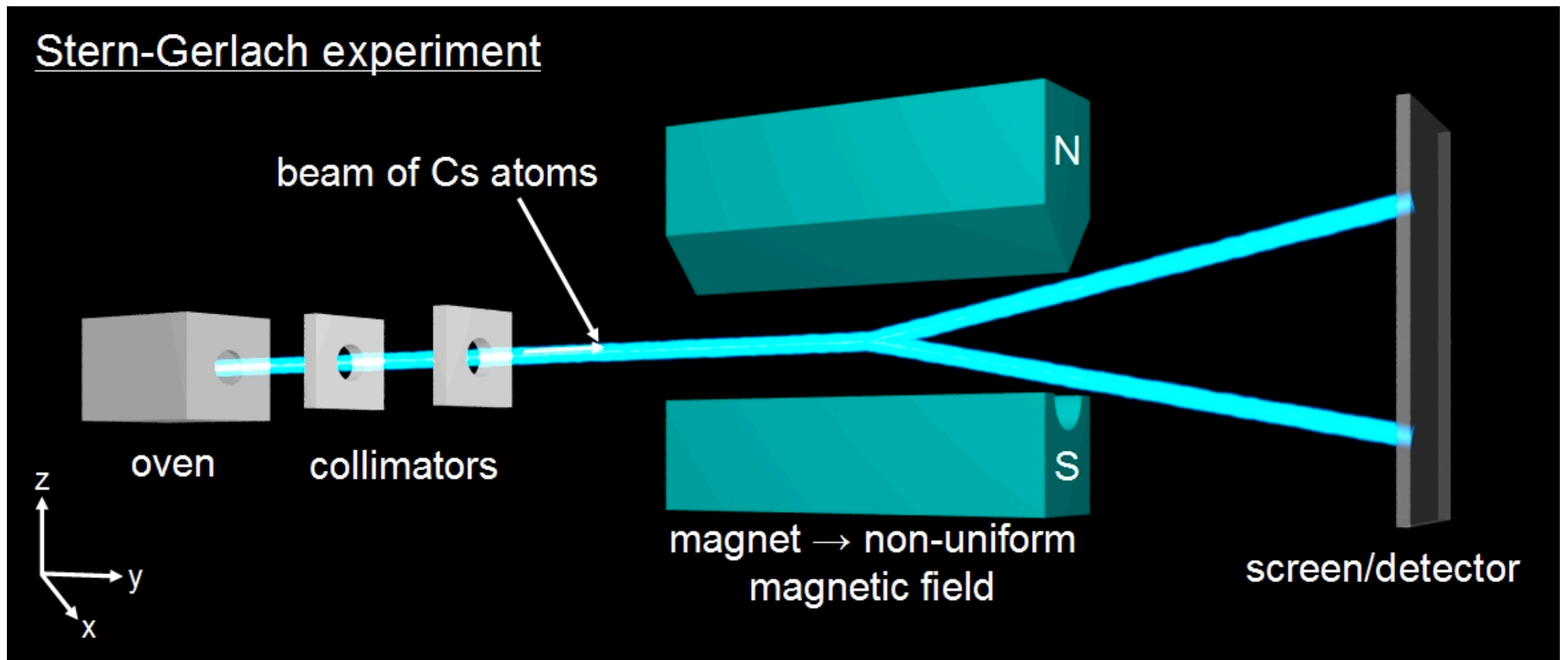
moment of inertia angular velocity

Intrinsic angular momentum

In classical mechanics, a solid body rotating around one axis has associated **angular momentum**, which is **conserved** in the absence of external forces

Around 1920, it was found that the electron appeared to have some **intrinsic angular momentum**, with only **two orientations** possible

$$|\vec{L}_e| = \frac{\hbar}{2} \quad L_{e,z} = \pm \frac{\hbar}{2}$$



exercise Intrinsic angular momentum

Can we understand intrinsic angular momentum of electron using **classical mechanics**?

classical mechanics

$$|\vec{L}_e| = I \times \omega = \left(\frac{2}{5} m_e R_e^2 \right) \omega$$

experiment

$$|\vec{L}_e| = \frac{\hbar}{2}$$

If we assume the electron is a solid (classical) body, **at which speed** it needs to rotate?

$$m_e = 9.11 \times 10^{-31} \text{ kg}$$

$$r_e \sim 10^{-18} \text{ m}$$

$$\hbar = 1.05 \times 10^{-34} \text{ J s}$$


exercise Intrinsic angular momentum

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$$\hbar = 1.05 \times 10^{-34} \text{ J s}$$

$$|\vec{L}_e| = \left(\frac{2}{5} m_e R_e^2 \right) \frac{v}{R_e} = \frac{\hbar}{2} \quad \rightarrow \quad v = \frac{5\hbar}{4m_e r_e} \simeq 1.5 \times 10^{14} \text{ m/s}$$

If we think of the intrinsic angular momentum of the electron in classical terms, its surface should move **faster than the speed of light**, which is impossible!

Spin

In quantum theory, we call **spin** the **intrinsic angular momentum** of elementary particles, which as we have just seen does not admit a classical interpretation

Spin is an example of a **quantum number**: intrinsic properties of elementary particles that might (electric charge) or not (color charge, spin) have classical counterparts

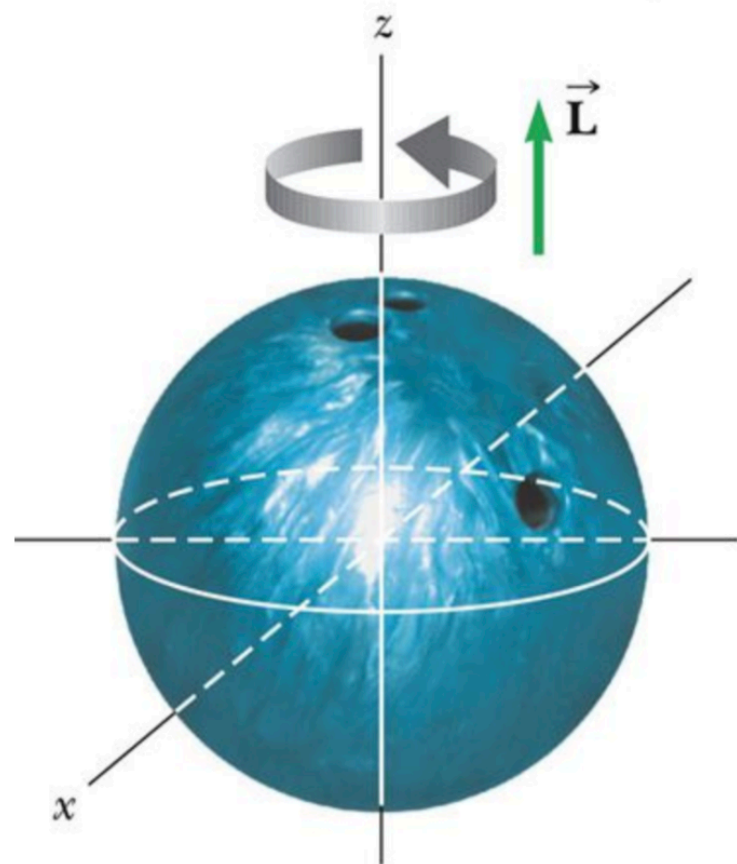
Spin

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Spin is quantised: it cannot take arbitrary values, but only a finite subset of them

Bowling ball (classical mechanics)



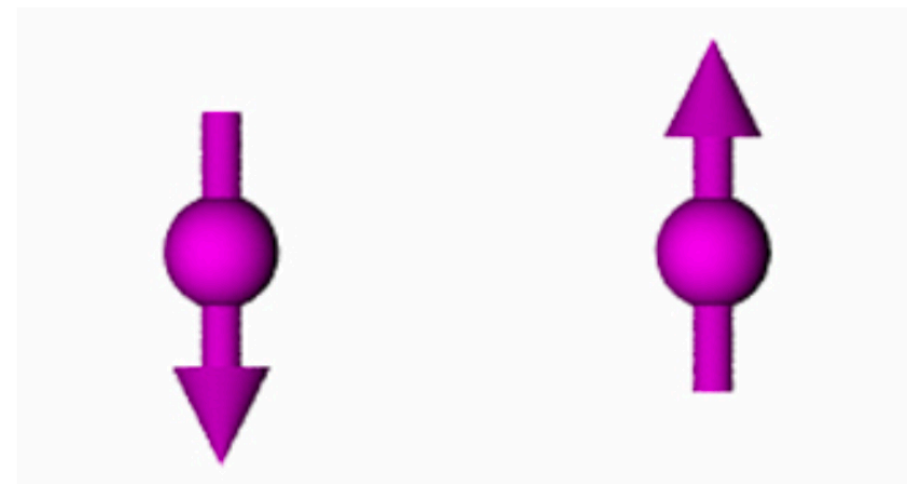
$$L = I\omega = \left(\frac{2}{5}MR\right)\omega$$

Juan Rojo

Electrons (quantum mechanics)

$|-\rangle$

$|+\rangle$



$$L = \frac{\hbar}{2}, L_z = -\frac{\hbar}{2}$$

$$L = \frac{\hbar}{2}, L_z = +\frac{\hbar}{2}$$

Introduction to Elementary Particles, 18/12/2018

Spin

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Electrons (quantum mechanics)

Electron with spin up

$$|\varphi\rangle_{\text{up}} = |+\rangle = |\uparrow\rangle$$

Electron with spin down

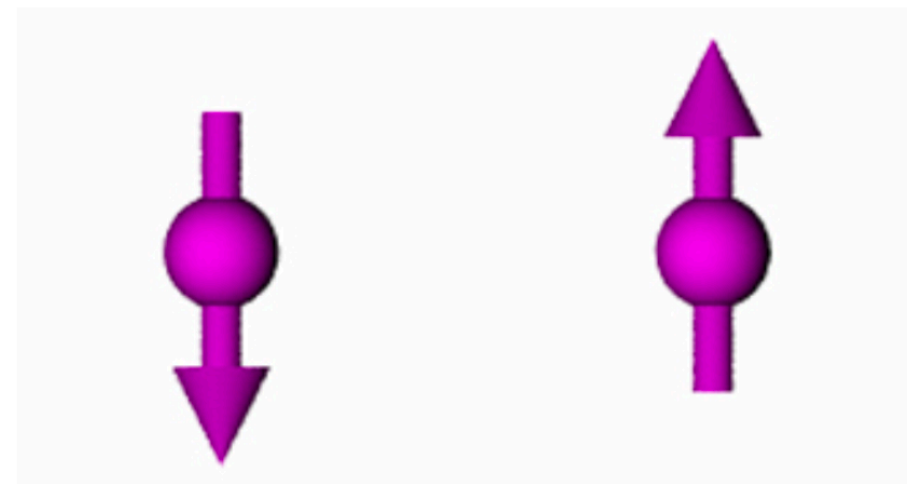
$$|\varphi\rangle_{\text{down}} = |-\rangle = |\downarrow\rangle$$

Electron with spin superposition

$$|\varphi\rangle_{\text{mixed}} = |+\rangle + |-\rangle = |\uparrow\rangle + |\downarrow\rangle$$

$|-\rangle$

$|+\rangle$



$$L = \frac{\hbar}{2}, L_z = -\frac{\hbar}{2}$$

$$L = \frac{\hbar}{2}, L_z = +\frac{\hbar}{2}$$

Fermions and bosons

Fermions and bosons

In terms of their **spin quantum number s** , particles can be divided into two groups:

Fermions: half-integer spin

$$s = \frac{\hbar}{2}, \frac{3\hbar}{2}, \dots$$

***Elementary
particles***

***Composite
particles***

Bosons: integer spin

$$s = 0, \hbar, 2\hbar, \dots$$

Fermions and bosons

In terms of their **spin quantum number s** , particles can be divided into two groups:

Fermions: half-integer spin

$$s = \frac{\hbar}{2}, \frac{3\hbar}{2}, \dots$$

Electron, muon, tauon,
neutrinos, quarks

$$s = \frac{\hbar}{2}$$

***Elementary
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***Composite
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Elementary particles

Electron, muon, tauon,
neutrinos, quarks $s = \frac{\hbar}{2}$

Composite particles

Protons, neutrons $s = \frac{\hbar}{2}$
Some **baryons**
e.g. Δ^{++} $s = \frac{3\hbar}{2}$

Bosons: integer spin

$$s = 0, \hbar, 2\hbar, \dots$$

Fermions and bosons

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Elementary particles

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Some **baryons**
e.g. Δ^{++} $s = \frac{3\hbar}{2}$

Bosons: integer spin

$$s = 0, \hbar, 2\hbar, \dots$$

Higgs boson $s = 0$
Photon, gluon, W , Z $s = \hbar$
Graviton (?) $s = 2\hbar$

Fermions and bosons

In terms of their **spin quantum number s** , particles can be divided into two groups:

Fermions: half-integer spin

$$s = \frac{\hbar}{2}, \frac{3\hbar}{2}, \dots$$

Elementary particles

Electron, muon, tauon, neutrinos, quarks

$$s = \frac{\hbar}{2}$$

Composite particles

Protons, neutrons

$$s = \frac{\hbar}{2}$$

Some **baryons**

e.g. Δ^{++}

$$s = \frac{3\hbar}{2}$$

Bosons: integer spin

$$s = 0, \hbar, 2\hbar, \dots$$

Higgs boson

$$s = 0$$

Photon, gluon, W , Z

$$s = \hbar$$

Graviton (?)

$$s = 2\hbar$$

Pions, kaons

$$s = 0$$

Some **mesons**

e.g. ρ , ω

$$s = 1$$

Fermions and bosons

In terms of their **spin quantum number s** , particles can be divided into two groups:

Fermions
Spin 1/2

Quarks:
up, down, strange,
charm, bottom, top

Charged leptons:
electrons, muons, tauons

Neutral leptons:
electron neutrino,
muon neutrino,
tau neutrino

Bosons

Spin 0
Higgs boson

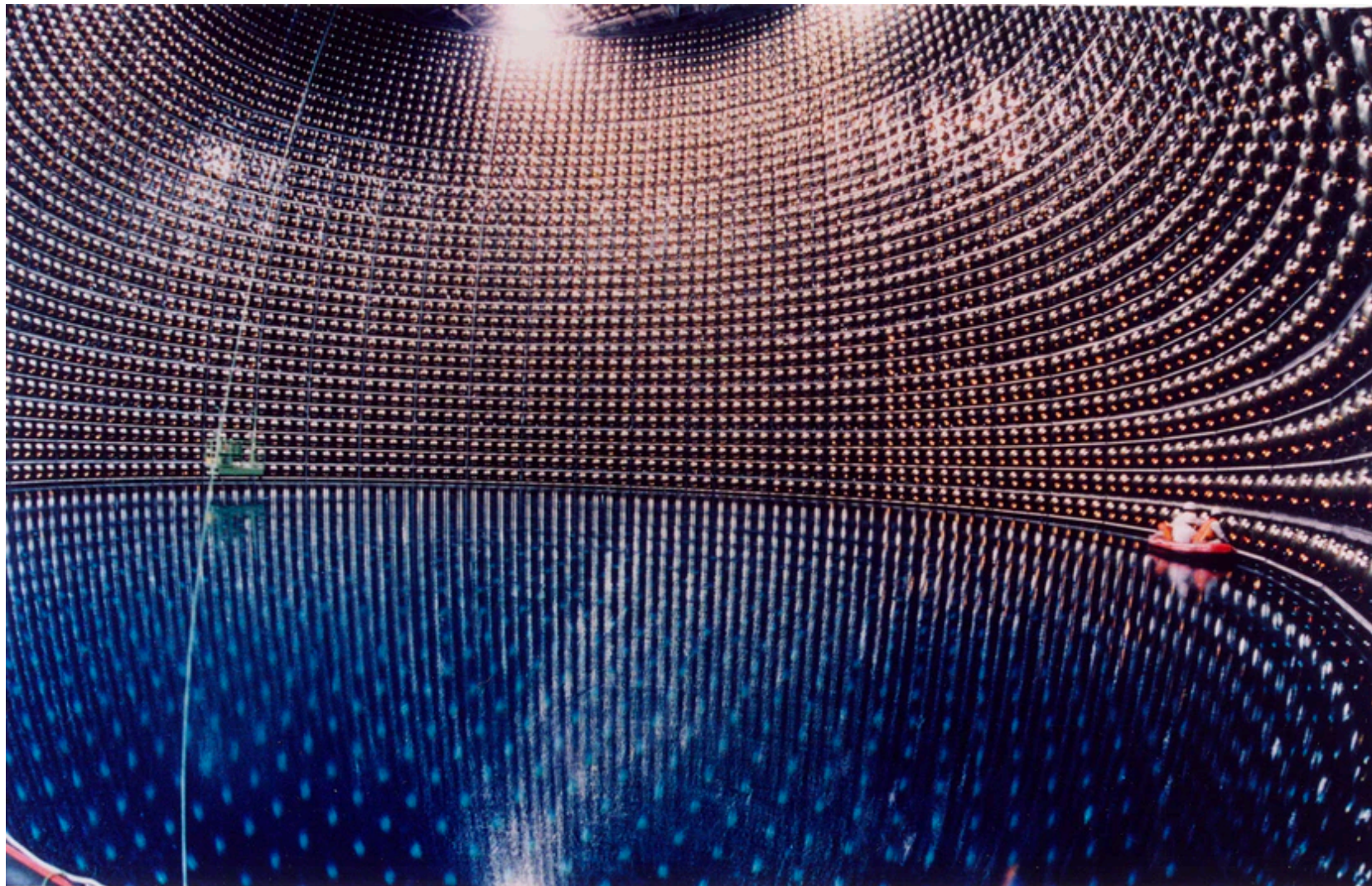
Spin 1
Photon (electromagnetism)
Gluon (strong force)
W,Z bosons (weak force)

Spin 2
Graviton (gravity)

**How large a neutrino
detector needs to be?**

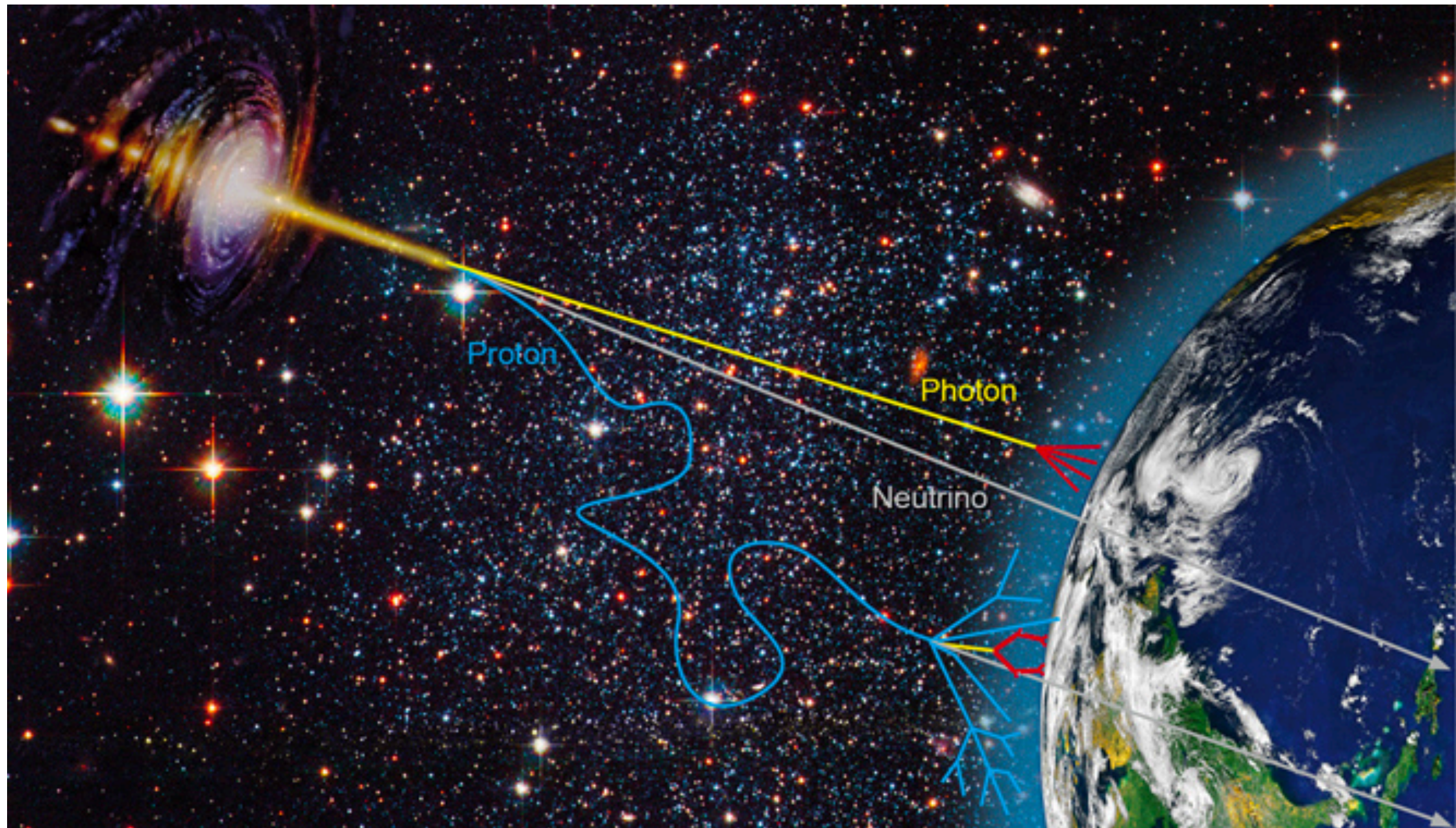
Neutrinos: the ghost particles

- ☑ **Neutrinos** are electrically neutral, very light, and weakly interacting particles introduced by Pauli in 1930 to guarantee **energy conservation** in the beta decay process
- ☑ Due to their very low interaction rate they are extremely difficult to measure, and huge detectors are required to be able to study them
- ☑ These ghostly particles are very abundant in the Universe: every second about **100 trillion neutrinos from the Sun cross your body!**



The neutrino universe

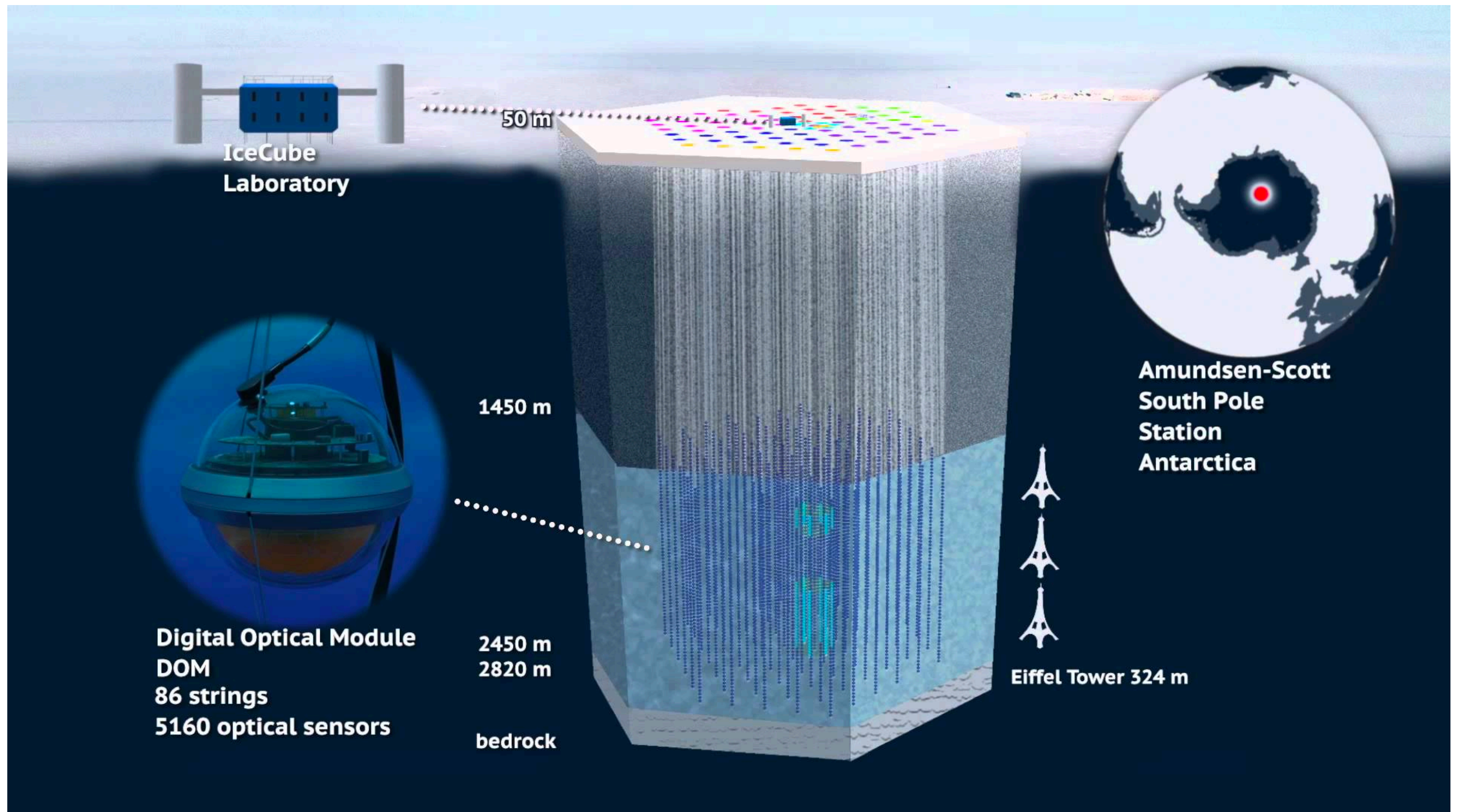
Neutrinos are not **deflected or attenuated**: unique probes of extreme astrophysical events



Neutrinos interact very weakly: need to **instrument huge volumes**, around **1 km³**, to have significant event rates

Neutrino telescopes

Ultra-high energy (UHE) neutrinos: novel window to the extreme Universe



Neutrino telescopes

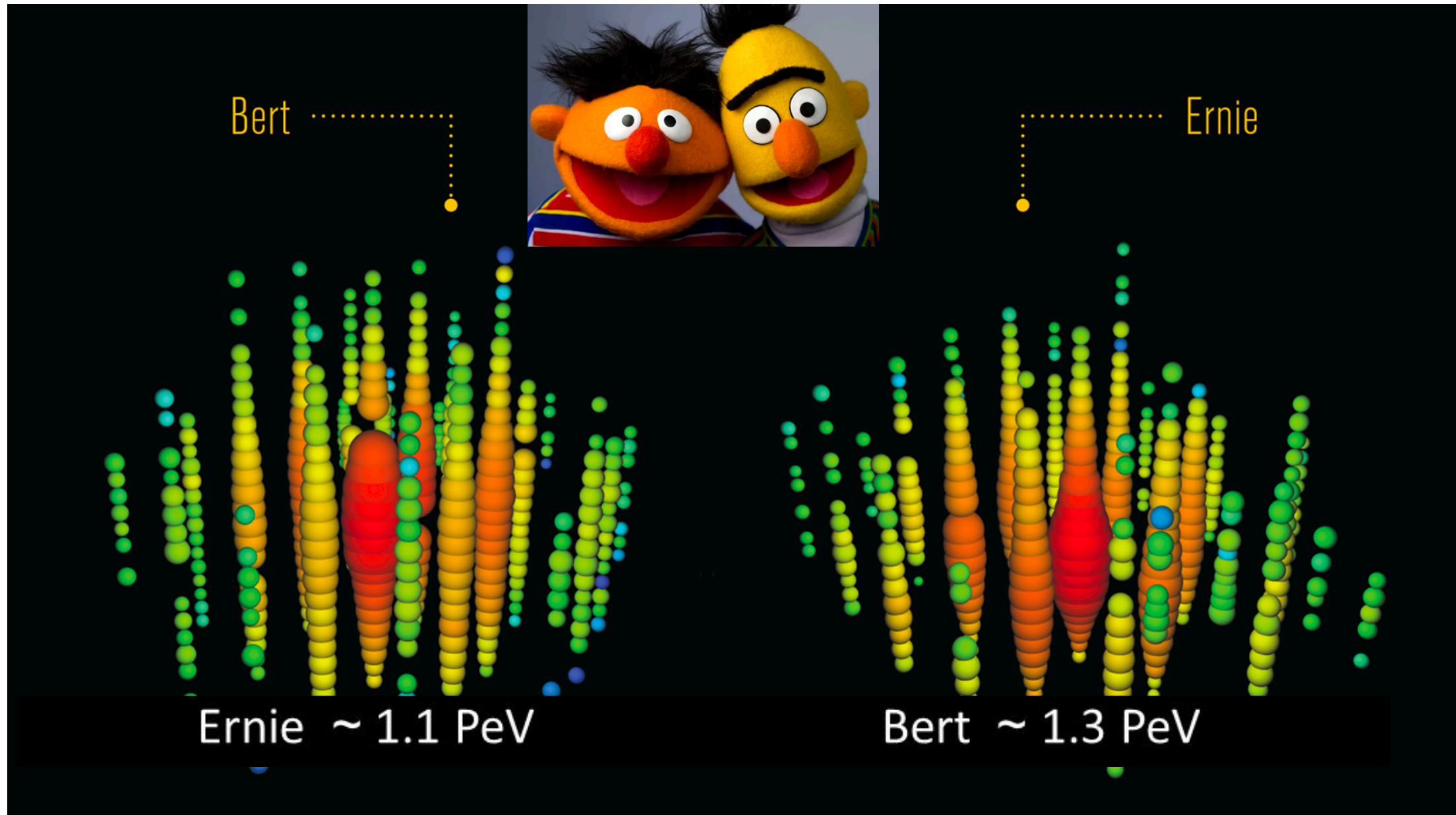
Neutrinos interact very weakly: need to **instrument huge volumes** (1 km^3) to have significant event rates

IceCube event
with simulated Cherenkov cone



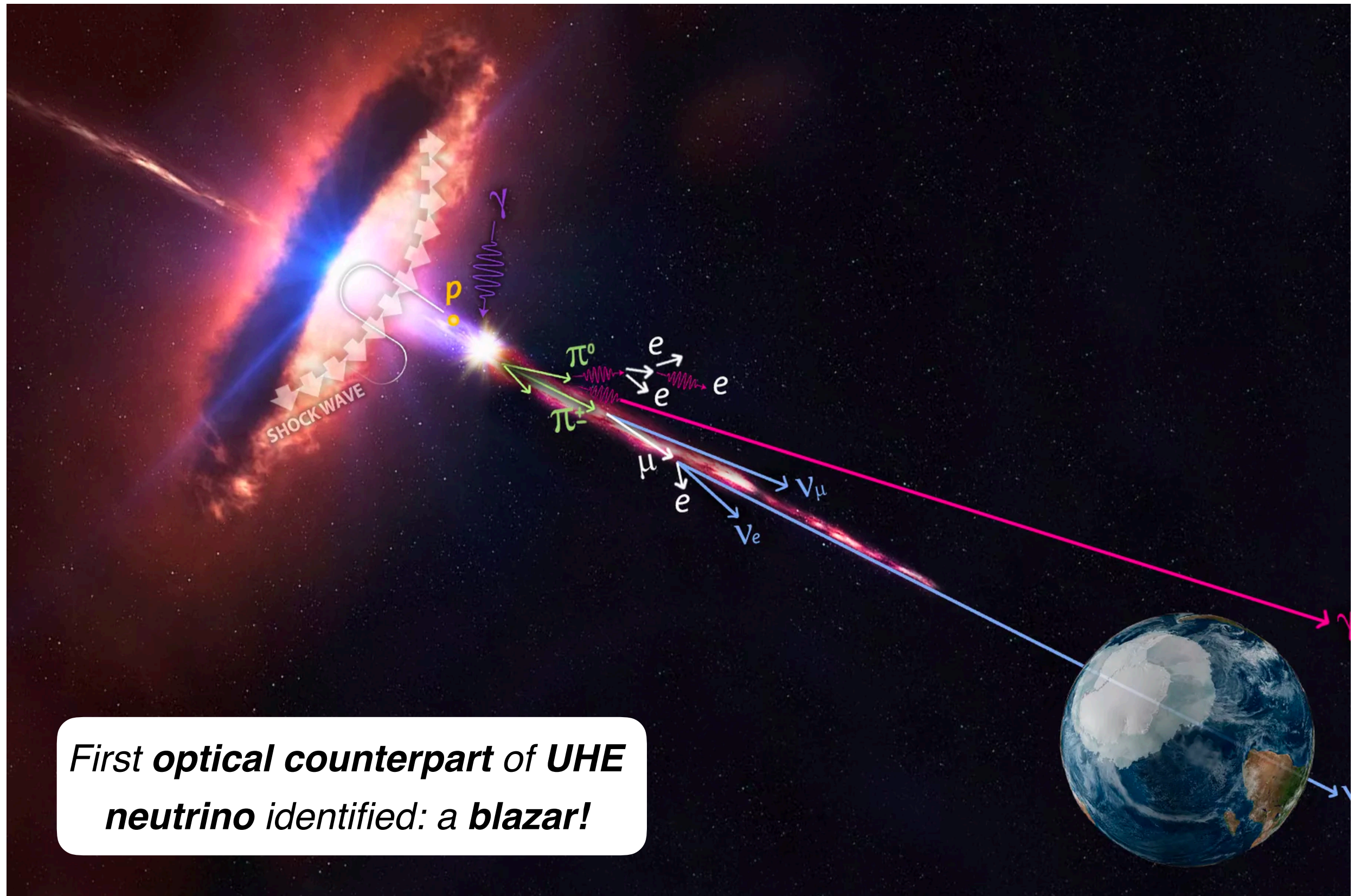
Neutrino telescopes

Neutrinos interact very weakly: need to **instrument huge volumes** (1 km^3) to have significant event rates



*Same or higher centre-of-mass energy than in **LHC collisions***

Unveiling cosmic neutrino origin

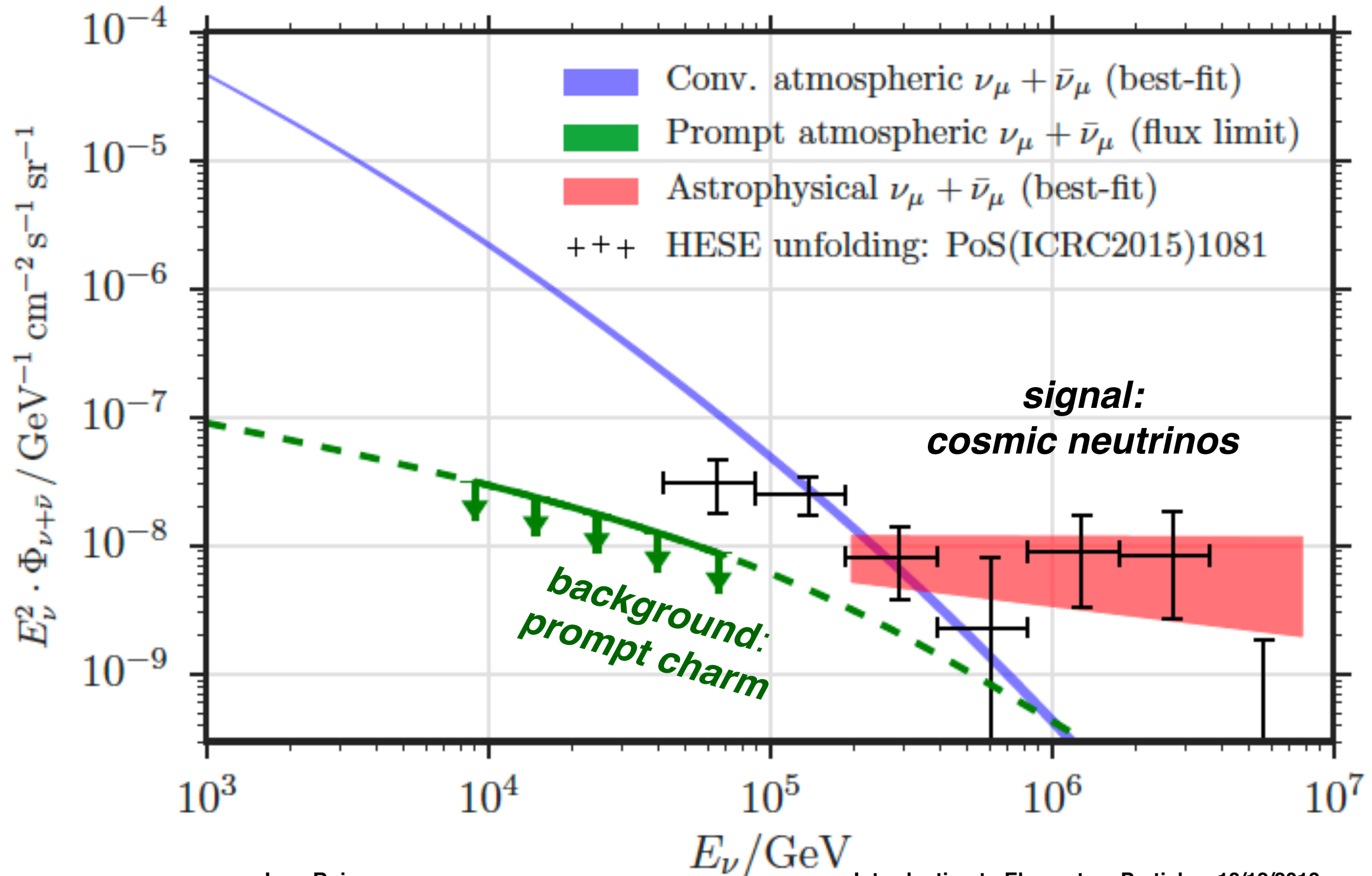


First optical counterpart of UHE neutrino identified: a blazar!

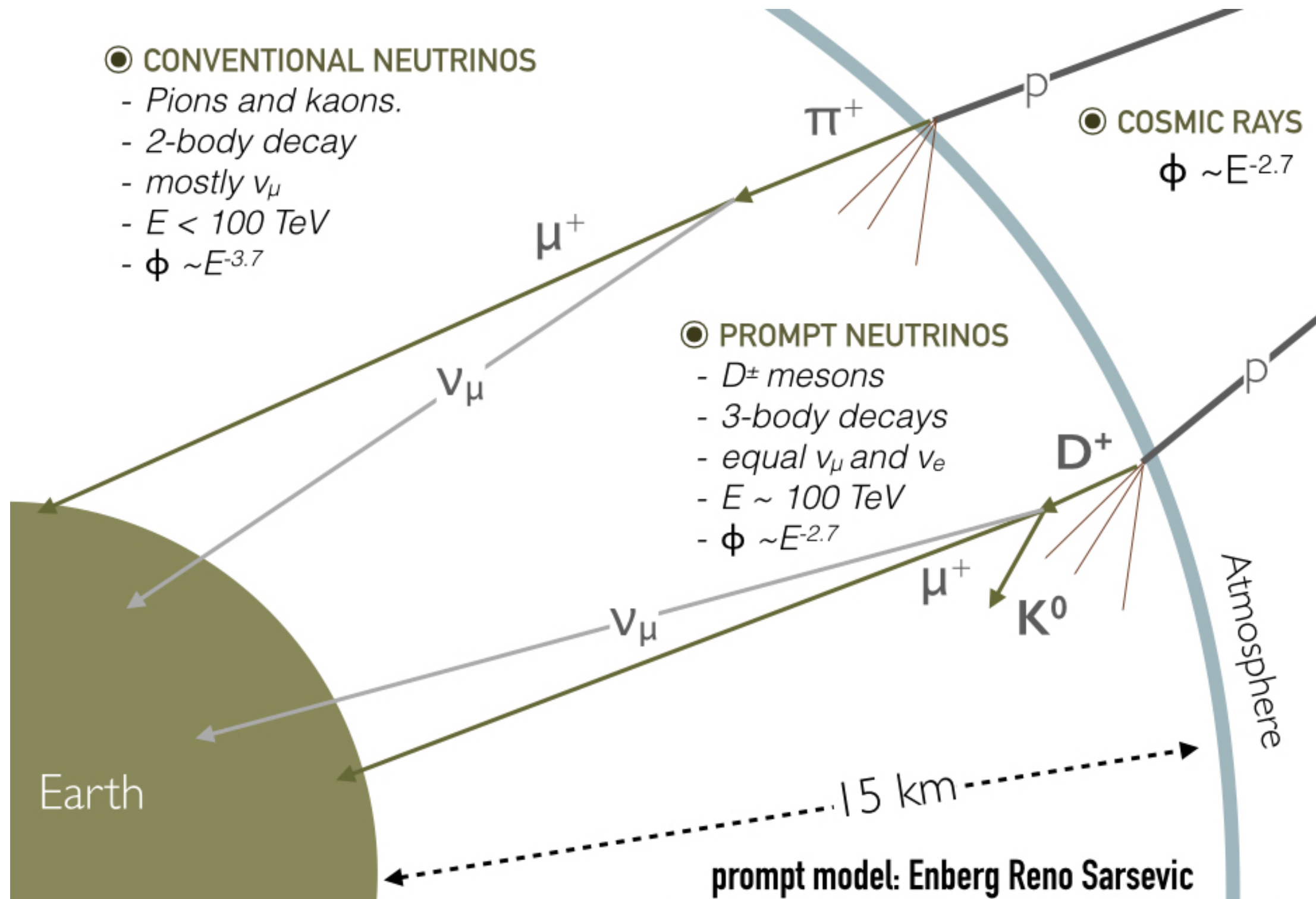
Neutrino telescopes as particle microscopes

signal: cosmic neutrino - nucleus scattering

background: prompt charm production



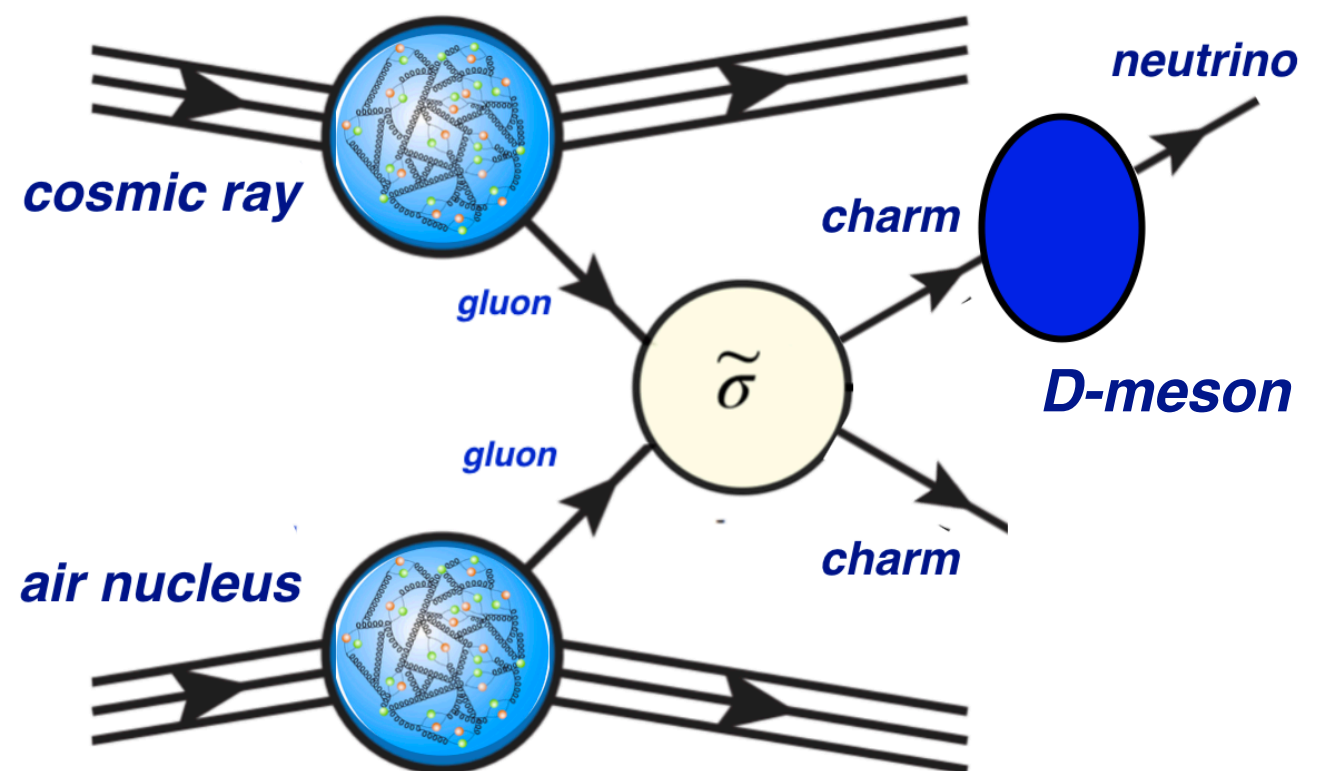
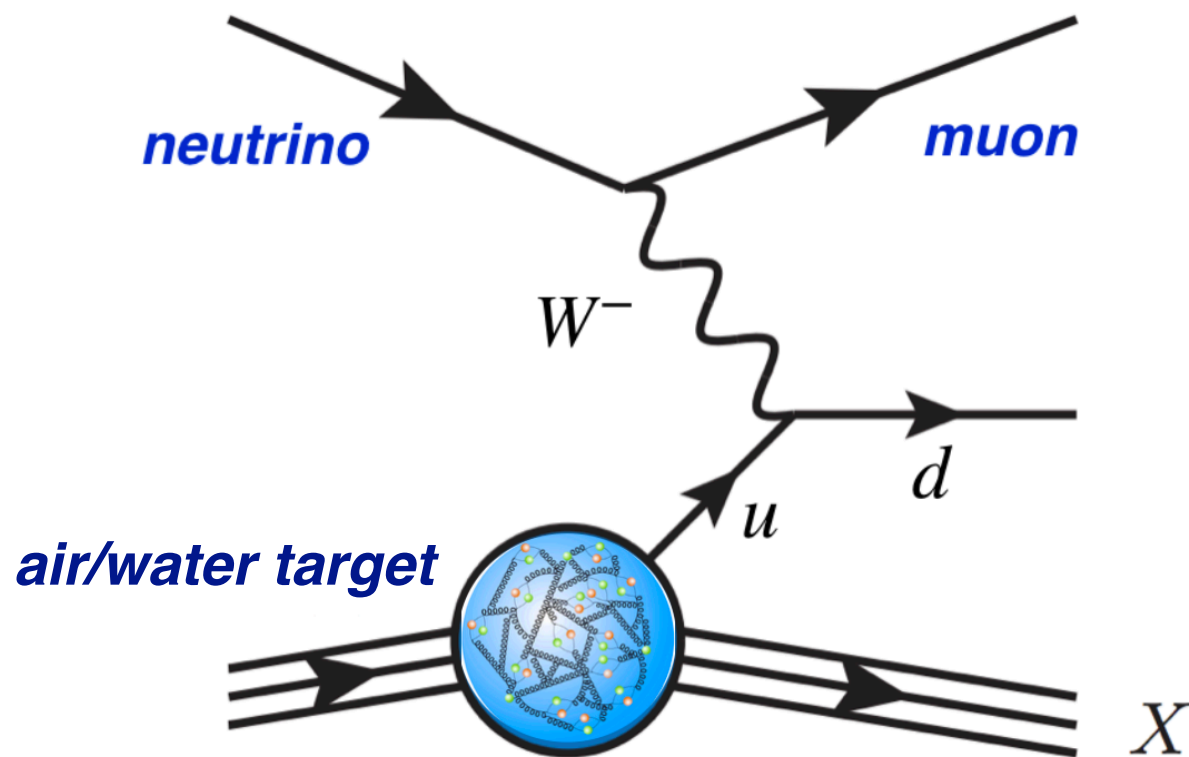
Neutrino telescopes as particle microscopes



Neutrino telescopes as particle microscopes

signal: cosmic neutrino - nucleus scattering

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High-energy particle astrophysics: unique connection between elementary particle physics, astronomy, and cosmology

exercise

Why 1 km³ for UHE neutrinos?

The **interaction cross-section** σ measures how likely a given scattering reaction is to take place. It is a kind of **effective collision area** and the units are cm⁻²

In general, the number of scattering events is given by

$$N_{\text{coll}} = \mathcal{L} \times \sigma \times T$$

where T is the exposure time, and the instantaneous **luminosity** \mathbf{L} is given by

$$\mathcal{L} = J \times n_t \times \Delta x$$

Flux of incoming particles

Target number density

Target thickness

Compute the **thickness of an ice target** if we want to detect at least **10 neutrinos per year** with energy of **10⁵ GeV**, using the following data

$$\sigma(\nu N, E_\nu = 10^5 \text{ GeV}) = 2 \times 10^{-34} \text{ cm}^2$$

$$n_t = 5.5 \times 10^{24} \text{ cm}^{-3}$$

$$J(E_\nu = 10^5 \text{ GeV}) = 3 \times 10^{-3} \text{ s}^{-1}$$

exercise

Why 1 km³ for UHE neutrinos?

First of all, compute the **instantaneous luminosity** as a function of the thickness

$$\mathcal{L} = 3 \times 10^{-3} \text{ s}^{-1} \times 5.5 \times 10^{24} \text{ cm}^{-3} \times \Delta x = (1.6 \times 10^{22} \times \Delta x) \text{ cm}^{-3} \text{ s}^{-1}$$

Then evaluate the **number of scattering collisions** expected during one year

$$N_{\text{coll}} = \mathcal{L} \times \sigma \times T = (1.6 \times 10^{22} \times \Delta x) \text{ cm}^{-3} \text{ s}^{-1} \times 2 \times 10^{-34} \text{ cm}^2 \times 3 \times 10^7 \text{ s}$$

Finally demand this number to be greater than 10 collisions per year

$$N_{\text{coll}} \sim 10^{-4} \text{ cm}^{-1} \times \Delta x > 10$$

$$\Delta x > 10^5 \text{ cm} \sim 1 \text{ km}$$

Therefore we need an **ice target of 1 km** to be able to record enough UHE events

This is size instrumented by **IceCube** (south pole) and **KM3NET** (Mediterranean sea)

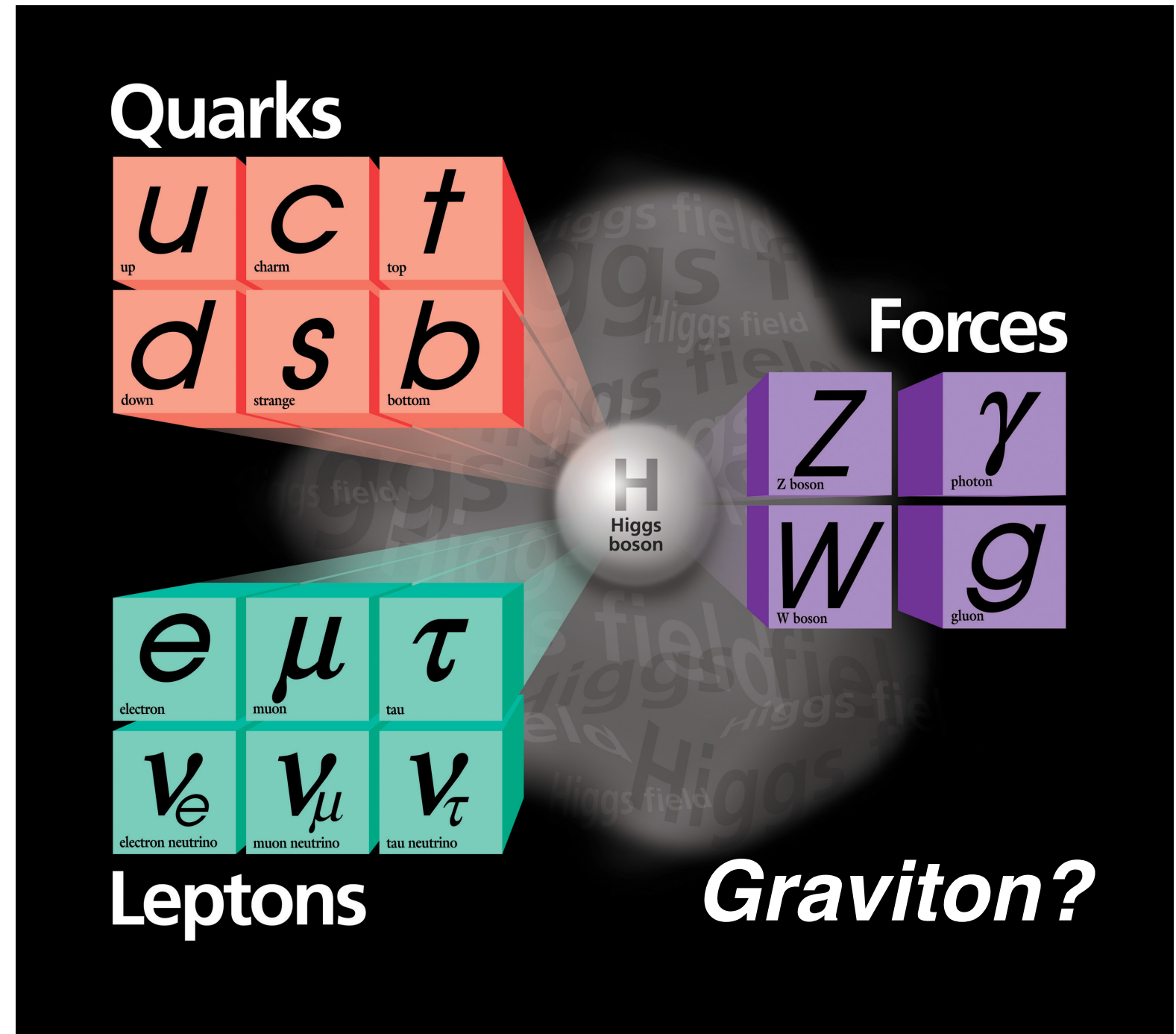
Gravity and gravitational waves

The gravitational interaction

☑ The Standard Model does not include the **gravitational interaction**

☑ **Gravity** is described by Einstein's theory of **General Relativity**

☑ So far we have not been able to construct a **quantum theory of gravity**



exercise

Is gravity strong or weak?

- ☑ Of the four fundamental interactions, **gravity's presence** is the one most familiar in everyday life
- ☑ Since we notice gravity more than other forces, does it mean that **gravity is much stronger**?

Compare the **gravitational** and **electric forces** between a **proton** and an **electron** within an hydrogen atom: which interaction is stronger?

$$F_g = G \frac{m_p m_e}{r^2} \qquad F_e = k \frac{q_p q_e}{r^2}$$

using the following data:

$$G = 6.7 \times 10^{-11} \text{ N m}^2 \text{ kg}^{-2}$$

$$k = 9 \times 10^9 \text{ N m}^2 \text{ C}^{-2}$$

$$m_e = 9.11 \times 10^{-31} \text{ kg}$$

$$q_e = -q_p = -1.6 \times 10^{-19} \text{ C}$$

$$m_p = 1.67 \times 10^{-27} \text{ kg}$$

exercise

Is gravity strong or weak?

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- ☑ Since we notice gravity more than other forces, does it mean that **gravity is much stronger?**

Compare the **gravitational** and **electric forces** between a **proton** and an **electron** within an hydrogen atom: which interaction is stronger?

$$F_g = G \frac{m_p m_e}{r^2} \qquad F_e = k \frac{q_p q_e}{r^2}$$

$$\frac{F_g}{F_e} = 10^{-39}$$

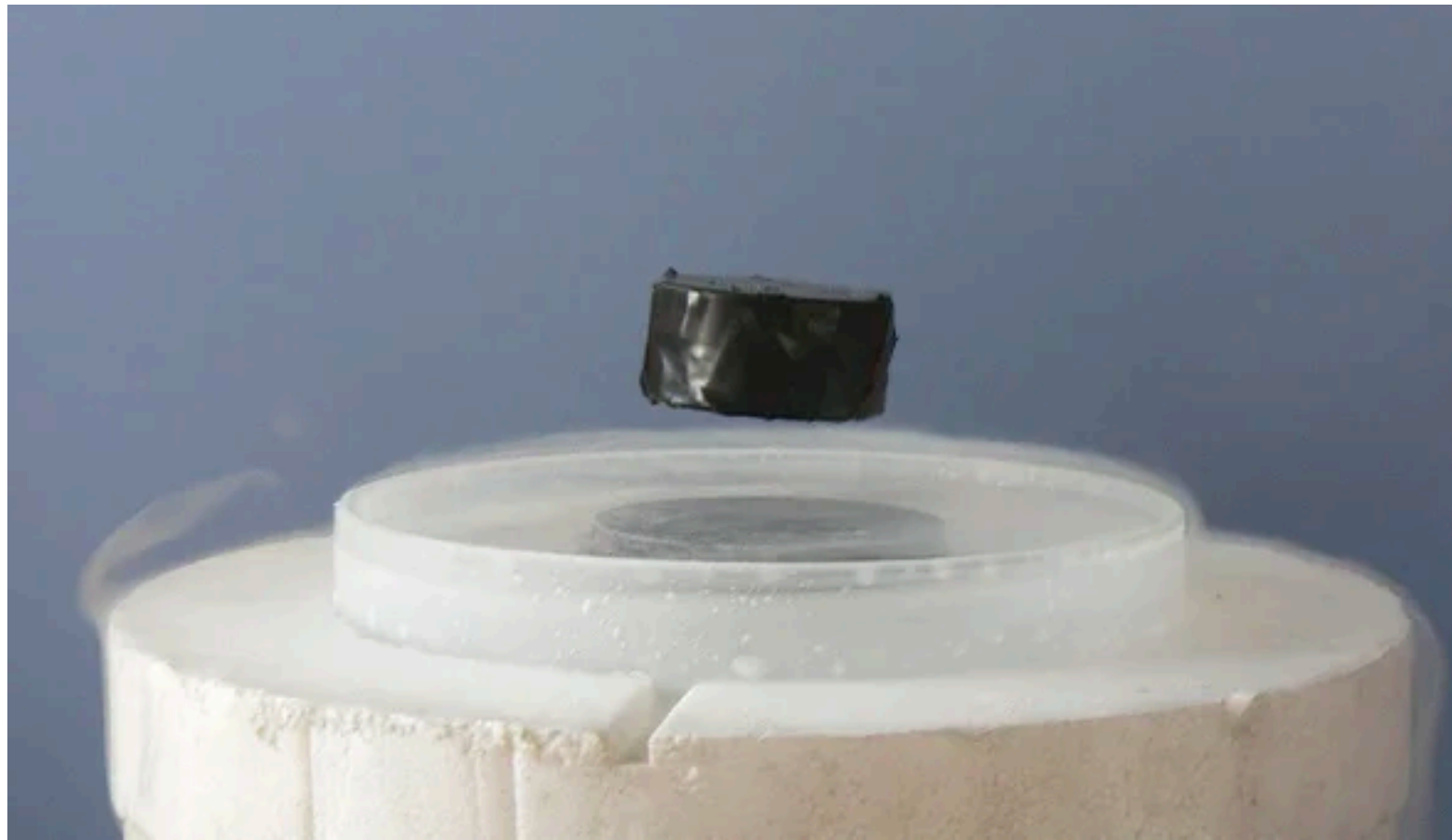
Gravity is **immensely weaker** than electromagnetism!

exercise

Is gravity strong or weak?

A small magnetic field can lift a magnet, **overcoming the gravitational pull** of the whole planet Earth

Why then **appears** that gravity is much stronger than the electric force?



Gravity is **immensely weaker** than electromagnetism!

exercise

Is gravity strong or weak?

A small magnetic field can lift a magnet, **overcoming the gravitational pull** of the whole planet Earth

Why then **appears** that gravity is much stronger than the electric force?

Gravity

All masses are *positive*:
gravitational interaction
always attractive

Electromagnetism

Electric charges can be positive and negative, thus electric interaction either **attractive or repulsive**: large bodies electrically neutral

Gravity is **immensely weaker** than electromagnetism!

Ripples in space-time



Ripples in space-time



Einstein's theory of General Relativity describes **gravity** as **deformations of space-time**
Massive enough bodies, such as black holes, deform space-time and **slow the local time**

Black holes

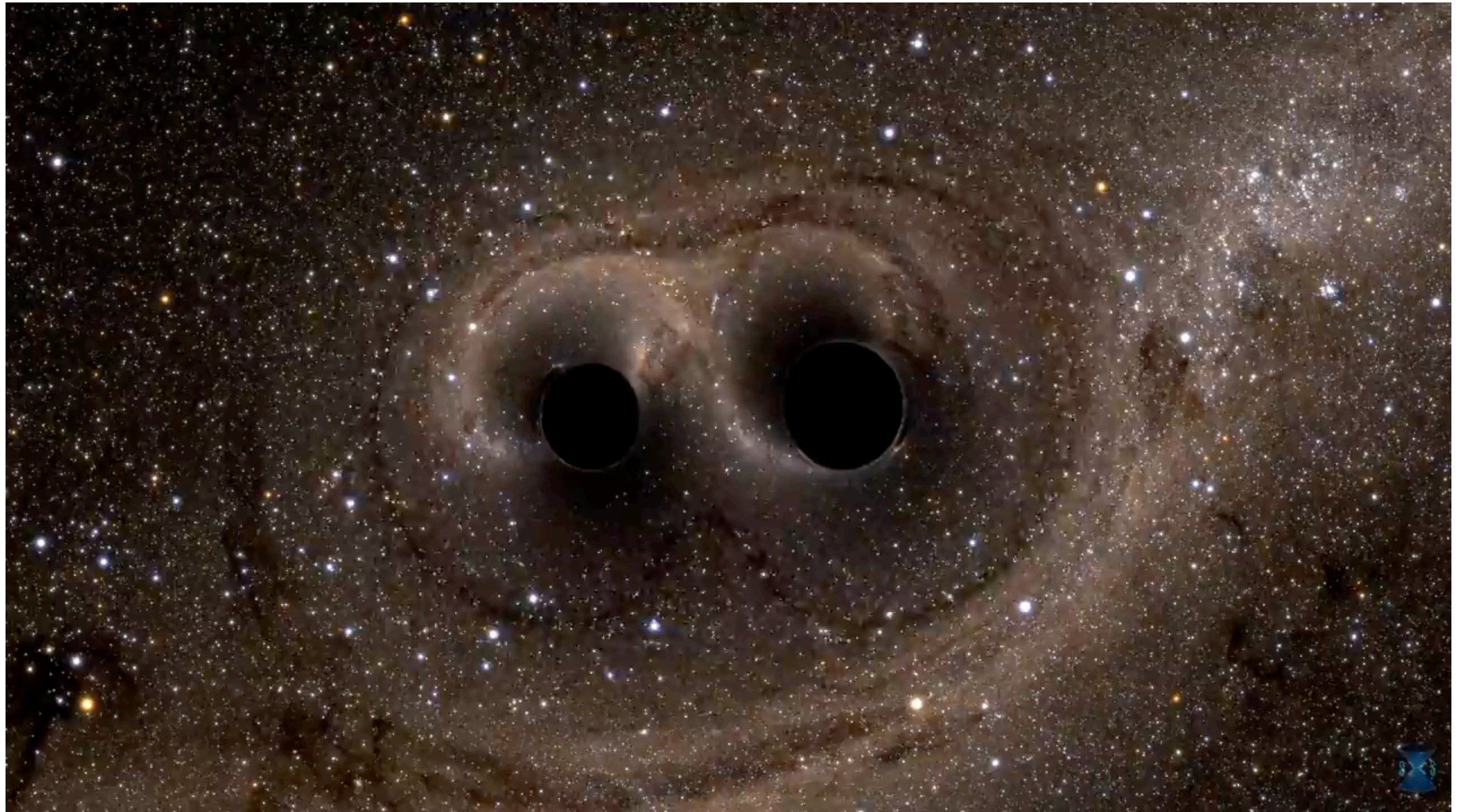
Black holes are one of the most fascinating objects in the Universe:
nothing can escape from their attraction, not even light!



When worlds collide

Black holes can be found in **pairs**, remnants from binary star systems

Black hole mergers are some of the most cataclysmic events that take place in the Universe

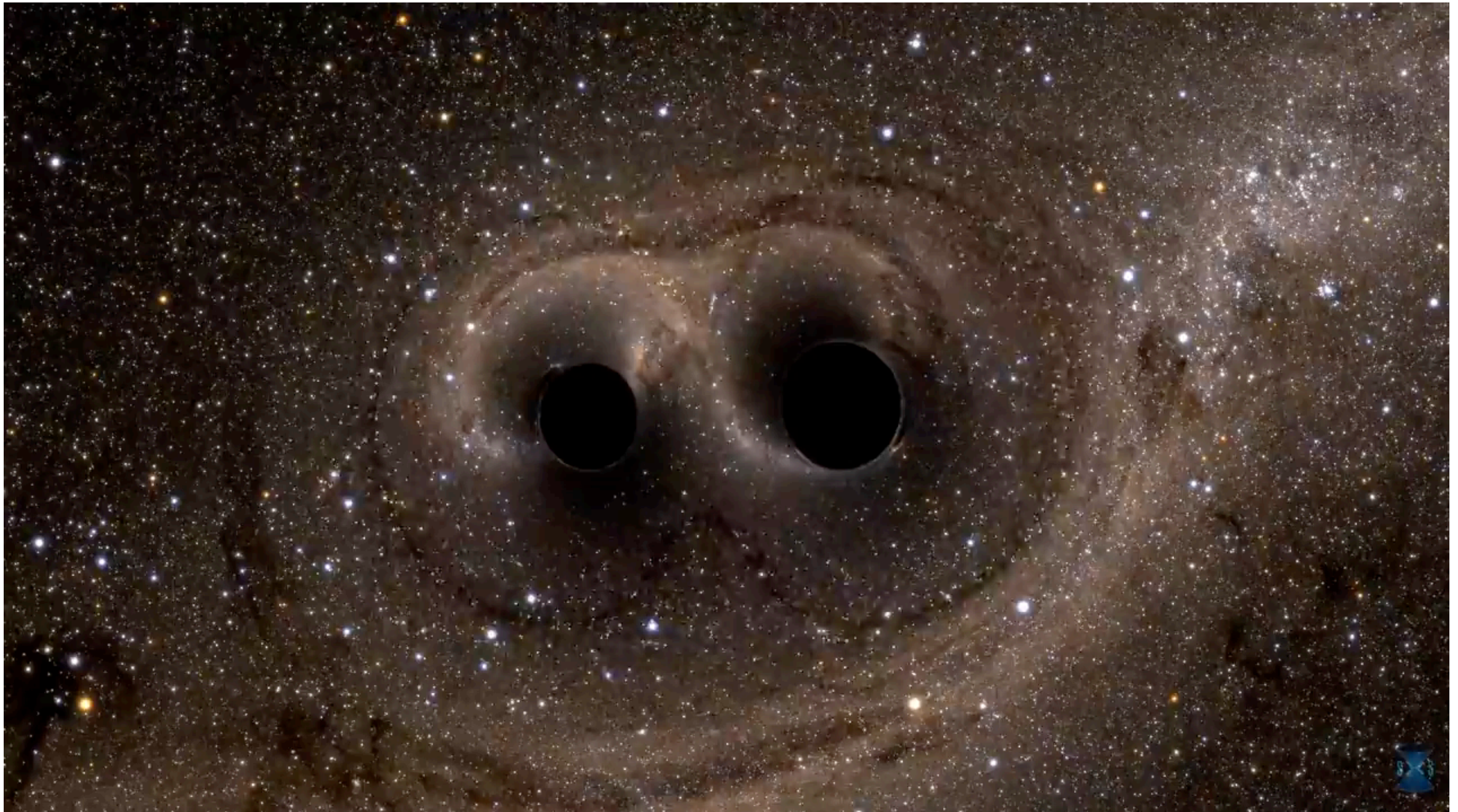


If a black hole is, well, black, then how is all this energy released?

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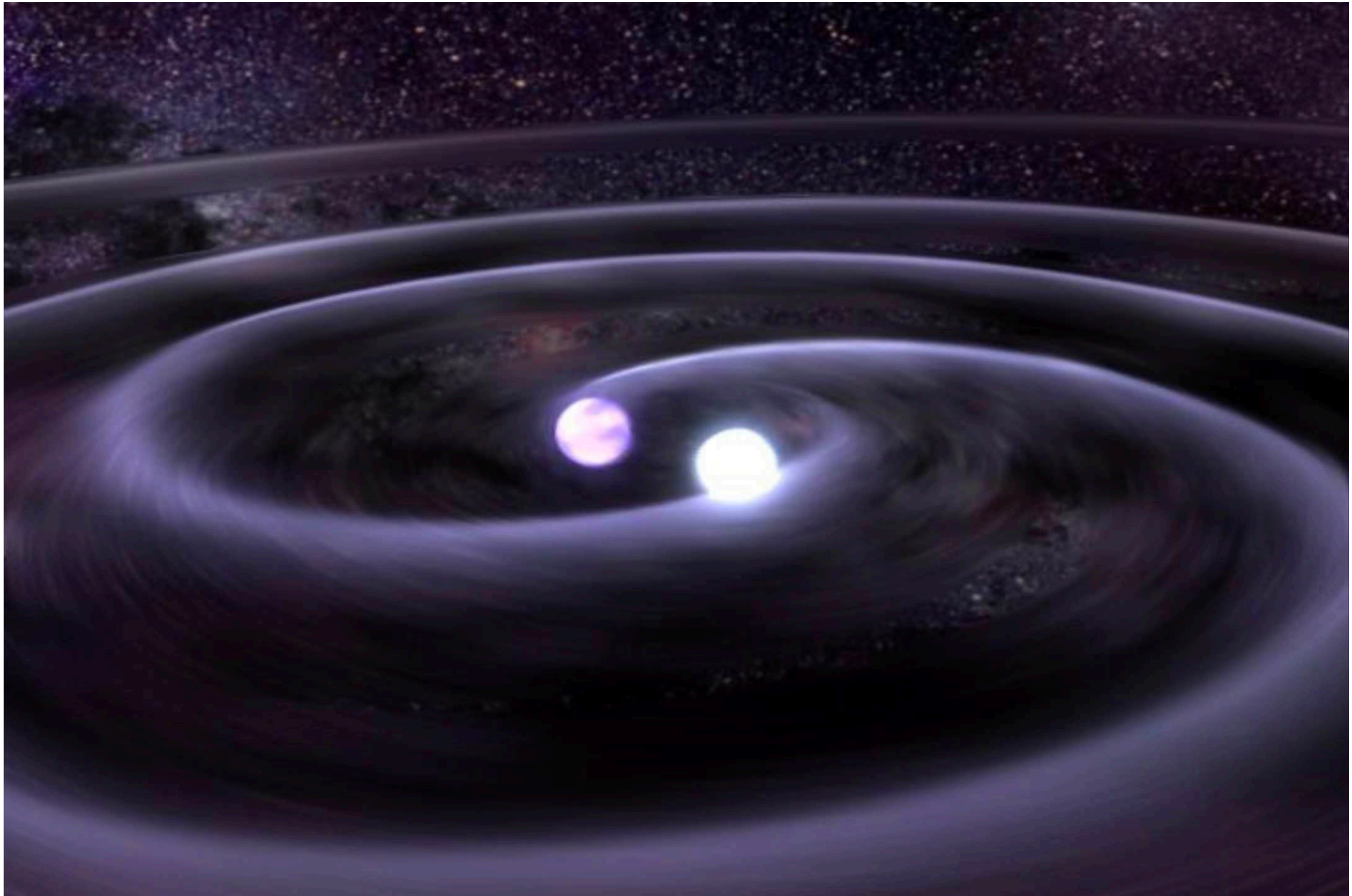


If a black hole is, well, black, then how is all this energy released?

In terms of gravitational waves, ripples of space-time itself!

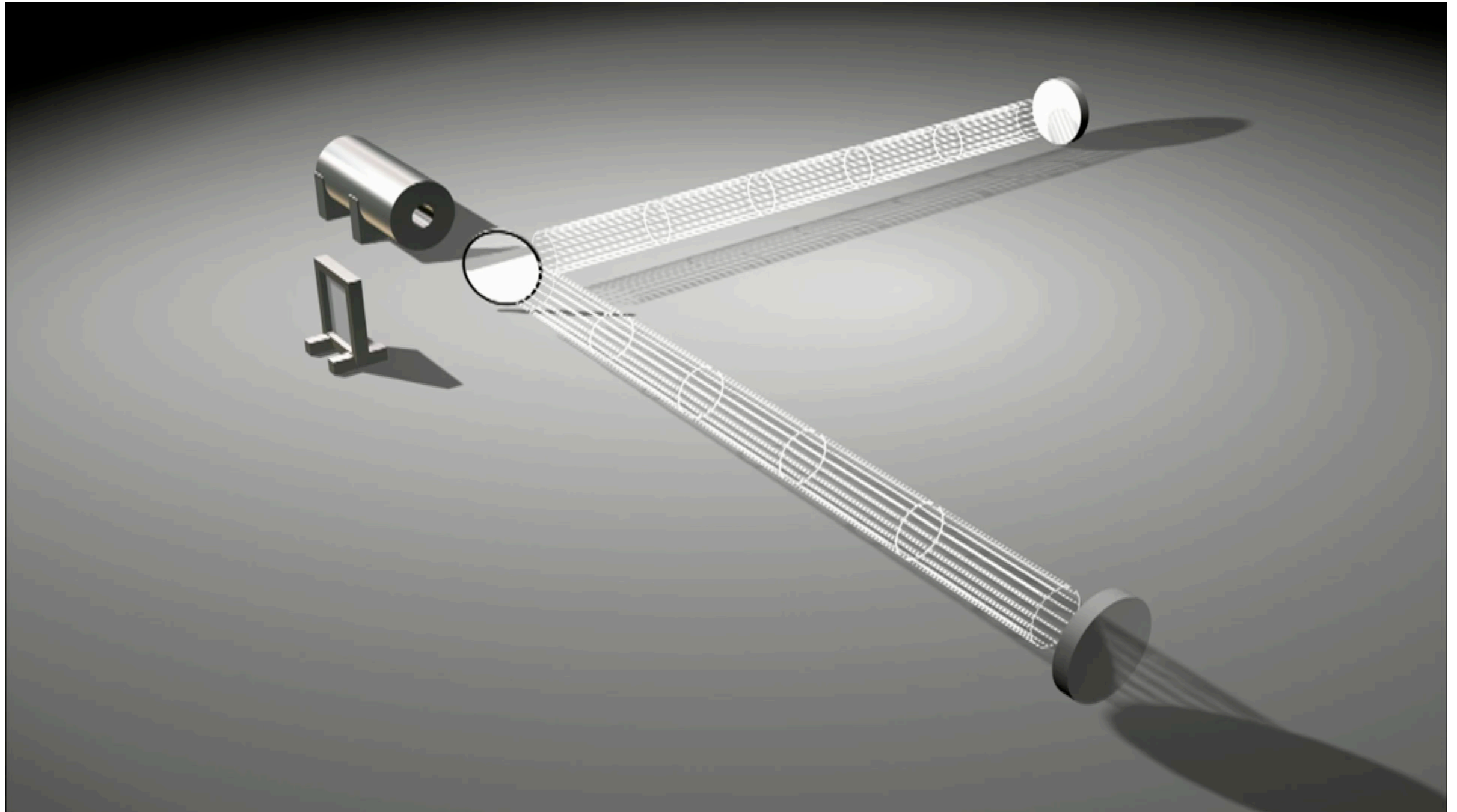
Ripples in space time

*Gravitational waves, unlike matter waves, do not propagate on top of something
space-time itself oscillates, propagating energy across the universe*



Detecting gravitational waves

Gravitational waves can be detected with **ultra-precise laser interferometers**



Need to measure **length variations of less than 1/1000** of a proton size in the interferometer arms of 4 km each!

Detecting gravitational waves

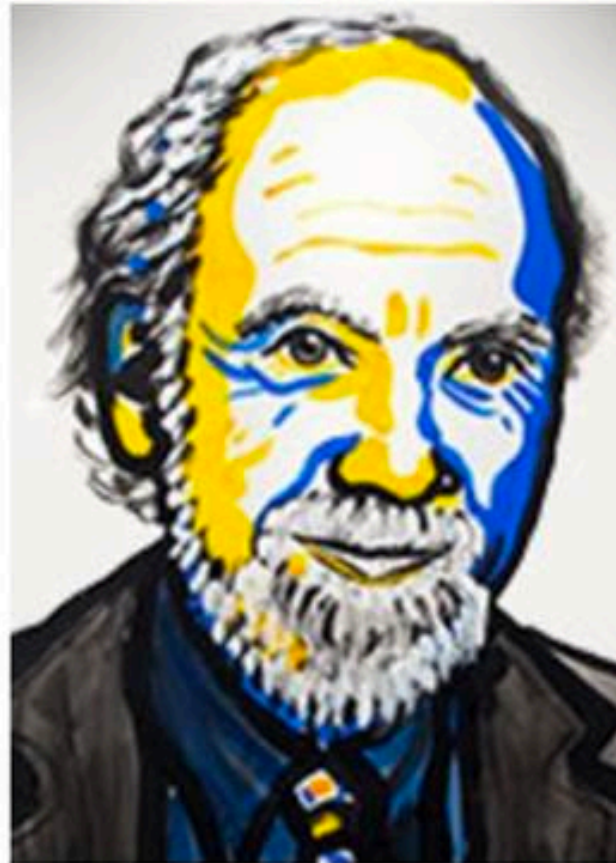
Nobel Prize in Physics 2017!



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Rainer Weiss

Prize share: 1/2



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Barry C. Barish

Prize share: 1/4



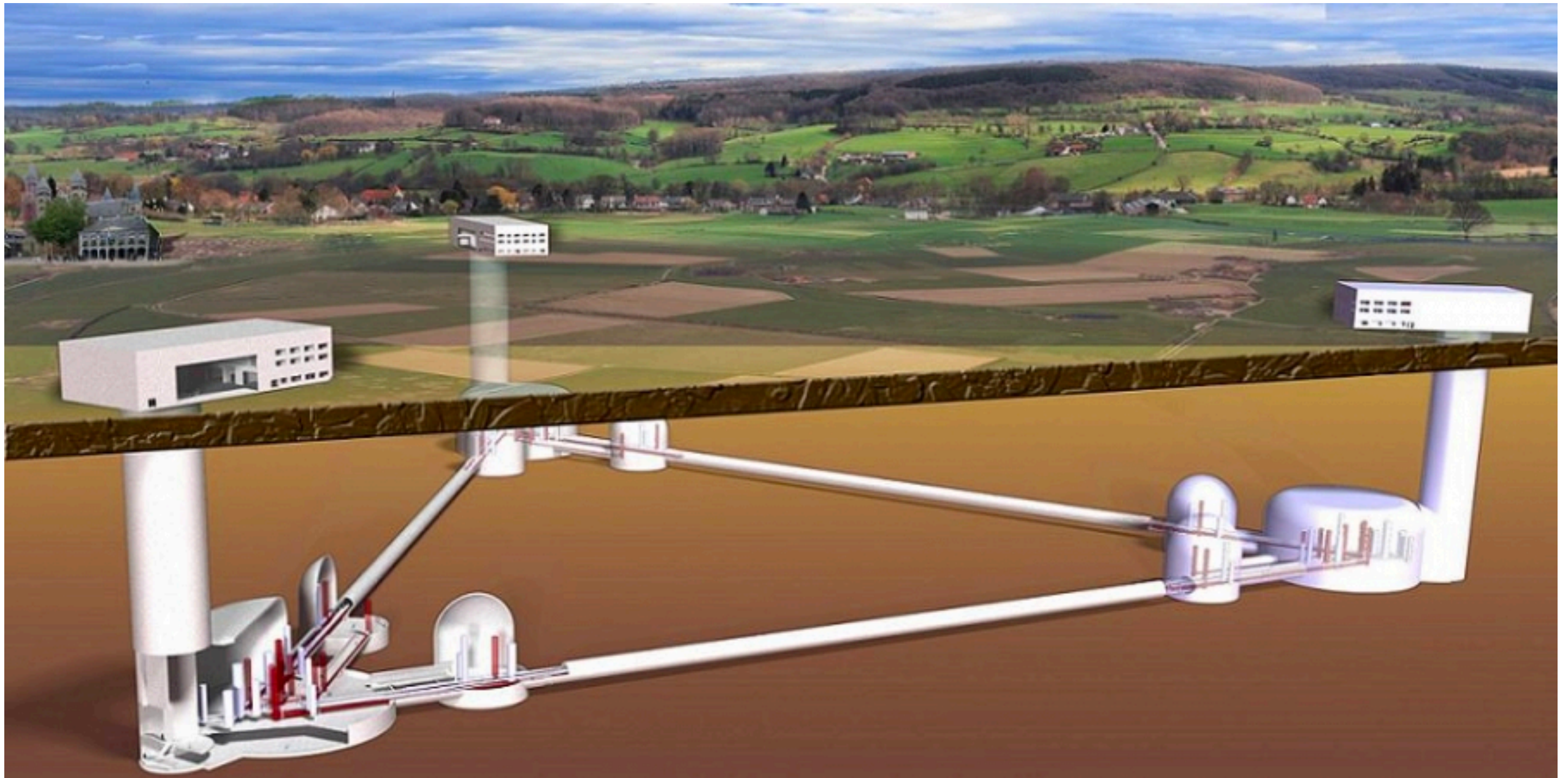
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Kip S. Thorne

Prize share: 1/4

Detecting gravitational waves

Strong **Dutch** involvement in Gravitational Wave detection!



Einstein Telescope: a next-generation gravitational wave observatory in Limburg

Special Relativity for elementary particles

Four-momenta

In Special Relativity **energy** and **linear momenta** are reference-frame dependent

Their description when changing reference frames is facilitated using **four-momenta**

$$p^\mu = \left(E/c, p_x, p_y, p_z \right) \quad \mu = 0, 1, 2, 3$$

The **norm** of a four-vector is the **same in all reference frames**

$$A^\mu = (A^0, A^1, A^2, A^3) \rightarrow |A| = \left((A^0)^2 - (A^1)^2 - (A^2)^2 - (A^3)^2 \right)^{1/2}$$

Note the differences as compared to the **Euclidean form** (Pythagoras Theorem)

For a particle with **mass m** moving at a **speed v** (in 1D), then E and p_x given by

$$E = \left(1 - v^2/c^2 \right)^{-1/2} mc^2$$

$$p_x = \left(1 - v^2/c^2 \right)^{-1/2} mv$$

exercise

Four-momenta

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Compute the norm of its four-momentum. Why it makes sense that this norm is exactly the same in every reference frame?

Four-momenta

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$$p_x = \left(1 - v^2/c^2\right)^{-1/2} mv$$

Compute the norm of its four-momentum. Why it makes sense that this norm is exactly the same in every reference frame?

$$p^\mu = (E/c, p_x, 0, 0) = \left(1 - v^2/c^2\right)^{-1/2} (mc, mv, 0, 0)$$

$$|p^\mu| = \left(1 - v^2/c^2\right)^{-1/2} \left(m^2c^2 - m^2v^2\right)^{1/2} = mc$$

For a single particle, the **norm of its four-momentum is its mass**, which of course should be the same in all reference frame

Time dilation

In special relativity, **time intervals** depend on the observer's reference frame

Assume that in a given reference frame some event lasts for a **period Δt** , as measured by a **co-moving clock**

An observer moving with velocity **u** with respect to this reference frame will measure a different time interval **$\Delta t'$** for exactly the same event

$$\Delta t' = \gamma \Delta t = \left(1 - u^2/c^2\right)^{-1/2} \Delta t > \Delta t$$

hence this phenomenon is called **time dilation**

Time dilation also affects elementary particles: for example, a particle with a given **half-life in its rest-frame** will have a **different half-life** in the lab frame

$$\tau|_{\text{lab}} = \left(1 - u^2/c^2\right)^{-1/2} \tau|_{\text{rest}}$$

exercise

Time dilation

$$\tau|_{\text{lab}} = \left(1 - u^2/c^2\right)^{-1/2} \tau|_{\text{rest}}$$

Example: assume that you have produced a pion moving at $\mathbf{v=0.9983c}$ in the laboratory frame. Evaluate the average decay length of this pion in the lab.

$$\tau|_{\text{rest}} = 2.6 \times 10^{-8} \text{ s}$$

exercise

Time dilation

$$\tau|_{\text{lab}} = \left(1 - u^2/c^2\right)^{-1/2} \tau|_{\text{rest}}$$

Example: assume that you have produced a pion moving at $v=0.9983c$ in the laboratory frame. Evaluate the average decay length of this pion in the lab.

$$\tau|_{\text{rest}} = 2.6 \times 10^{-8} \text{ s}$$

Its **average decay length** will be given by

$$\Delta x = v \times \tau|_{\text{lab}} = v \times \gamma \times \tau|_{\text{rest}} \simeq 70 \text{ m}$$

Without accounting for time dilation, you would have found that the pion should decay on average after 7 m, so your **detector would be all wrong!**

Summary and next steps

- 📌 We have presented a general overview of the world of **elementary particles and interactions**
- 📌 The theory of elementary particles provides a mathematical language allowing the description of a wide variety of phenomena, from the **structure of nucleons** to **particle colliders** and **high-energy astrophysics**
- 📌 After the Christmas break, we will present this language and to explore systematically the **basic properties of elementary particles and interactions**, which determine the outcome of scattering processes
- 📌 Please review the basic concepts of **Special Relativity** and **Quantum Theory** (see appendices in lecture notes) since in the following we will use them extensively